



Introduction to density-functional theory

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European Summerschool in Quantum Chemistry (ESQC) September 2022, Sicily, Italy

www.lct.jussieu.fr/pagesperso/toulouse/presentations/presentation_esqc_22.pdf

Why and how learning density-functional theory?

Density-functional theory (DFT) is:

- ▶ a practical electronic-structure computational method, widely used in quantum chemistry and condensed-matter physics;
- ▶ an exact and elegant reformulation of the quantum many-body problem, which has led to new ways of thinking in the field.

Classical books:

- R. G. Parr and W. Yang, Density-Functional Theory of Atoms and Molecules, Oxford University Press, 1989.
 R. M. Dreizler and E. K. U. Gross, Density Functional Theory: An Approach to the
- Quantum Many-Body Problem, Springer-Verlag, 1990.
 ▶ W. Koch and M. C. Holthausen, A Chemist's Guide To Density Functional Theory,
- My lecture notes:

Wiley-VCH, 2001.

http://www.lct.jussieu.fr/pagesperso/toulouse/enseignement/introduction_dft.pdf

A book chapter:

J. Toulouse, in *Density Functional Theory*, edited by E. Cancès and G. Friesecke (Springer), to appear; https://arxiv.org/abs/2103.02645

Outline

- Basic density-functional theory
 - Quantum many-electron problem
 - Universal density functional
 - Kohn-Sham method
 - Generalized Kohn-Sham method
- 2 Exact constraints for the exchange-correlation functional
 - Exact expressions for the exchange and correlation functionals
 - Uniform coordinate scaling
 - One-orbital spatial regions and self-interaction
 - Lieb-Oxford lower bound
 - Frontier orbital energies
- Usual approximations for the exchange-correlation energy
 - Local-density approximation
 - Semilocal approximations
 - Single-determinant hybrid approximations
 - Multideterminant hybrid approximations
 - Dispersion corrections
- 4 Additional topics in density-functional theory
 - Time-dependent density-functional theory
 - Some less usual orbital-dependent exchange-correlation functionals
 - Fractional electron numbers and frontier orbital energies

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 Levy-Lieb constrained-search formulation
 Density-matrix or Lieb density functional
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 The Kohn-Sham equations
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 Extension to spin density-functional theory
 - Generalized Kohn-Sham method

Quantum many-electron problem

- We consider an N-electron system in the Born-Oppenheimer and non-relativistic approximations.
- ► The **electronic Hamiltonian** in the position representation is, in atomic units,

$$H = -rac{1}{2}\sum_{i=1}^{N}
abla_{\mathbf{r}_{i}}^{2} + rac{1}{2}\sum_{i=1}^{N}\sum_{\substack{j=1\i = i}}^{N}rac{1}{|\mathbf{r}_{i} - \mathbf{r}_{j}|} + \sum_{i=1}^{N}
u_{ ext{ne}}(\mathbf{r}_{i})$$

 $H\Psi(x_1, x_2, ..., x_N) = E\Psi(x_1, x_2, ..., x_N)$

- where $v_{ne}(\mathbf{r}_i) = -\sum_{\alpha} Z_{\alpha}/|\mathbf{r}_i \mathbf{R}_{\alpha}|$ is the nuclei-electron interaction potential.
- ► Stationary states satisfy the time-independent Schrödinger equation

where
$$\Psi(\mathbf{x}_1, \mathbf{x}_2, ..., \mathbf{x}_N)$$
 is a wave function written with space-spin coordinates

- $\mathbf{x}_i = (\mathbf{r}_i, \sigma_i)$ (with $\mathbf{r}_i \in \mathbb{R}^3$ and $\sigma_i \in \{\uparrow, \downarrow\}$) which is antisymmetric with respect to the exchange of two space-spin coordinates, and E is the associated energy.
- ▶ We work on the **Hilbert space** $\mathcal{H} = \bigwedge^N L^2(\mathbb{R}^3 \times \{\uparrow, \downarrow\}, \mathbb{C}).$
- ▶ Using **Dirac notations** (representation-independent formalism):

$$\hat{H}|\Psi
angle=E|\Psi
angle \qquad ext{where} \qquad \hat{H}=\hat{T}+\hat{W}_{ ext{ee}}+\hat{V}_{ ext{ne}}$$

These operators can be conveniently expressed in (real-space) second quantization.

Wave-function variational principle

► The ground-state electronic energy E₀ can be expressed with the wave-function variational principle

$$E_0 = \min_{\Psi \in \mathcal{W}} \langle \Psi | \hat{H} | \Psi
angle$$

where the minimization is done over the space of admissible normalized N-electron wave functions $\mathcal{W}=\{\Psi\in \bigwedge^N H^1(\mathbb{R}^3\times\{\uparrow,\downarrow\},\mathbb{C})\mid \langle\Psi|\Psi\rangle=1\}.$

Remark: If \hat{H} does not bind N electrons, then the minimum does not exist but the ground-state energy can still be defined as an infimum, i.e. $E_0 = \inf_{\Psi \in \mathcal{W}} \langle \Psi | \hat{H} | \Psi \rangle$.

 DFT is based on a reformulation of this variational theorem in terms of the one-electron density defined as

$$\textit{n}(\textbf{r}) = \textit{N} \int \cdots \int \left| \Psi(\textbf{x}, \textbf{x}_2, ..., \textbf{x}_N) \right|^2 d\sigma d\textbf{x}_2 ... d\textbf{x}_N$$

which is normalized to the electron number, $\int n(\mathbf{r})d\mathbf{r} = N$.

Remark: Integration over a spin coordinate σ means a sum over the two values of σ , i.e. $\int d\sigma = \sum_{\sigma \in \{\uparrow,\downarrow\}}.$

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The Hohenberg-Kohn theorem

- Consider an electronic system with an arbitrary external local potential $v(\mathbf{r})$ (that bounds N electrons) in place of $v_{ne}(\mathbf{r})$.
- The corresponding ground-state wave function Ψ (or one of them if there are several) can be obtained by solving the Schrödinger equation, from which an associated **ground-state density** $n(\mathbf{r})$ can be deduced. Therefore, one has the (possibly multi-valued) mapping:

$$v(\mathbf{r}) \longrightarrow n(\mathbf{r})$$

In 1964, Hohenberg and Kohn showed that this mapping can be inverted, i.e. a **ground-state density** $n(\mathbf{r})$ determines the potential $v(\mathbf{r})$ up to an arbitrary additive constant:

$$n(\mathbf{r}) \xrightarrow{\mathsf{Hohenberg-Kohn}} v(\mathbf{r}) + \mathsf{const}$$

Proof of the Hohenberg-Kohn theorem (1/2)

This is a two-step proof by contradiction.

Consider two local potentials differing by more than an additive constant:

$$v_1(\mathbf{r}) \neq v_2(\mathbf{r}) + \text{const}$$

We have two Hamiltonians:

$$\hat{H}_1 = \hat{T} + \hat{W}_{ee} + \hat{V}_1$$
 with a ground state $\hat{H}_1 | \Psi_1 \rangle = E_1 | \Psi_1 \rangle$ and ground-state density $n_1(\mathbf{r})$ $\hat{H}_2 = \hat{T} + \hat{W}_{ee} + \hat{V}_2$ with a ground state $\hat{H}_2 | \Psi_2 \rangle = E_2 | \Psi_2 \rangle$ and ground-state density $n_2(\mathbf{r})$

1 We first show that $\Psi_1 \neq \Psi_2$:

Assume
$$\Psi_1=\Psi_2=\Psi$$
. Then we have:
$$(\hat{H}_1-\hat{H}_2)|\Psi\rangle=(\hat{V}_1-\hat{V}_2)|\Psi\rangle=(E_1-E_2)|\Psi\rangle$$

or, in position representation,

$$\left(\sum_{i=1}^{N} [v_1(\mathbf{r}_i) - v_2(\mathbf{r}_i)]\right) \Psi(\mathbf{x}_1, \mathbf{x}_2, ..., \mathbf{x}_N) = (E_1 - E_2) \Psi(\mathbf{x}_1, \mathbf{x}_2, ..., \mathbf{x}_N)$$

If $\Psi(\mathbf{x}_1, \mathbf{x}_2, ..., \mathbf{x}_N) \neq 0$ for at least one fixed set of $(\sigma_1, \sigma_2, ..., \sigma_N)$ and "almost" all $(\mathbf{r}_1, \mathbf{r}_2, ..., \mathbf{r}_N)$, which is true for "reasonably well behaved potentials", then it implies that $v_1(\mathbf{r}) - v_2(\mathbf{r}) = \text{const}$, in contradiction with the initial hypothesis.

⇒ Intermediate conclusion: two local potentials differing by more than an additive constant cannot have a common ground-state wave function.

Proof of the Hohenberg-Kohn theorem (2/2)

We now show than $n_1 \neq n_2$:

Assume $n_1 = n_2 = n$. Then, by the variational theorem, we have:

$$E_1 = \langle \Psi_1 | \hat{H}_1 | \Psi_1 \rangle < \langle \Psi_2 | \hat{H}_1 | \Psi_2 \rangle = \langle \Psi_2 | \hat{H}_2 + \hat{V}_1 - \hat{V}_2 | \Psi_2 \rangle = E_2 + \int [v_1(\mathbf{r}) - v_2(\mathbf{r})] \, n(\mathbf{r}) d\mathbf{r}$$

The strict inequality comes from the fact that Ψ_2 cannot be a ground-state wave function of \hat{H}_1 , as shown in the first step of the proof.

So, we have shown

$$E_1 < E_2 + \int [v_1(\mathbf{r}) - v_2(\mathbf{r})] n(\mathbf{r}) d\mathbf{r}$$

Symmetrically, by exchanging the role of system 1 and 2, we have the strict inequality

$$E_2 < E_1 + \int [v_2(\mathbf{r}) - v_1(\mathbf{r})] \ n(\mathbf{r}) d\mathbf{r}$$

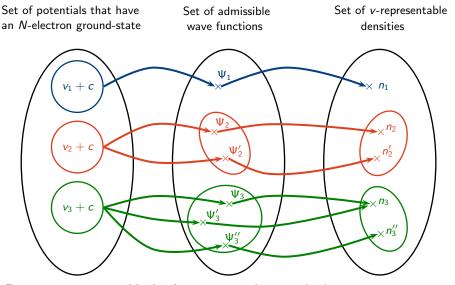
Adding the two inequalities gives the inconsistent result

$$E_1 + E_2 < E_2 + E_1$$

⇒ Conclusion: there cannot exist two local potentials differing by more than an additive constant which have a common ground-state density.

Remark: This proof does not assume non-degenerate ground states (contrary to the original Hohenberg-Kohn proof).

Hohenberg-Kohn theorem: Schematic illustration



From any v-representable density n, we can always go back to the potential v (up to an additive constant) having n as ground-state density.

The universal density functional and the variational property

▶ The Hohenberg-Kohn theorem can be summarized as $n \longrightarrow v \longrightarrow \hat{H} \longrightarrow \text{everything}$

$$n \longrightarrow V \longrightarrow H \longrightarrow \text{everyt}$$

v is a functional of the density n, i.e. v[n], and all other quantities as well. \triangleright For nondegenerate ground states, the ground-state wave function Ψ is a functional of

n, denoted by $\Psi[n]$. Hohenberg and Kohn defined the universal density functional $F[n] = \langle \Psi[n] | \hat{T} + \hat{W}_{ee} | \Psi[n] \rangle$

$$E[n] = F[n] + \int v_{ne}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r}$$

and the total electronic energy functional

Remark: In case of degeneracies, $\Psi[n]$ stands for any degenerate ground-state wave function

having the density n. Still, F[n] remains unique. ▶ The Hohenberg-Kohn universal functional is only defined for *N*-electron densities *n* that are ground-state densities associated with some local potential, the so-called set of *v*-representable densities which we will denote by A.

ground-state energy

Hohenberg and Kohn showed that we have a variational property giving the exact

The minimum is reached for an exact ground-state density $n_0(\mathbf{r})$ of the potential $v_{ne}(\mathbf{r})_{2/95}$

Levy-Lieb constrained-search formulation (1/2)

 In 1979 Levy, and later in 1983 Lieb, proposed to redefine the universal density functional as

$$\textit{F[n]} = \min_{\substack{\Psi \in \mathcal{W} \\ \Psi \rightarrow n}} \langle \Psi | \, \hat{\mathcal{T}} + \hat{\mathcal{W}}_{\text{ee}} | \Psi \rangle = \langle \Psi[n] | \, \hat{\mathcal{T}} + \hat{\mathcal{W}}_{\text{ee}} | \Psi[n] \rangle$$

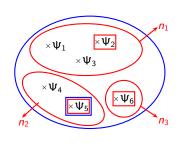
where " $\Psi \to n$ " means that the wave function Ψ is constrained to yield the fixed density n.

- ► This so-called Levy-Lieb functional *F*[*n*] does not require the existence of a local potential associated with the density.
- ▶ It is an extension of the Hohenberg-Kohn functional: it is defined on the larger set of N-electron densities coming from a wave function $\Psi \in \mathcal{W}$, the so-called set of N-representable densities $\mathcal{D} = \{n \in L^1(\mathbb{R}^3) \mid n \geq 0, \int n(\mathbf{r}) d\mathbf{r} = N, \sqrt{n} \in H^1(\mathbb{R}^3)\}.$

Levy-Lieb constrained-search formulation (2/2)

▶ The variational property is easily obtained using the constrained-search formulation:

$$\begin{split} E_0 &= \min_{\Psi \in \mathcal{W}} \langle \Psi | \, \hat{T} + \hat{W}_{ee} + \hat{V}_{ne} | \Psi \rangle \\ &= \min_{n \in \mathcal{D}} \min_{\substack{\Psi \in \mathcal{W} \\ \Psi \to n}} \langle \Psi | \, \hat{T} + \hat{W}_{ee} + \hat{V}_{ne} | \Psi \rangle \\ &= \min_{n \in \mathcal{D}} \left\{ \min_{\substack{\Psi \in \mathcal{W} \\ \Psi \to n}} \langle \Psi | \, \hat{T} + \hat{W}_{ee} | \Psi \rangle + \int v_{ne}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r} \right\} \\ &= \min_{n \in \mathcal{D}} \left\{ F[n] + \int v_{ne}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r} \right\} \end{split}$$



Hence, in DFT, we replace "min" by "min" which is a **tremendous simplification**! However, $F[n] = T[n] + W_{\text{ee}}[n]$ is **very difficult to approximate**, in particular the kinetic energy part T[n].

Density-matrix or Lieb density functional

In 1980 Valone and in 1983 Lieb generalized the constrained-search approach from pure states to ensemble density matrices Γ̂, resulting in the density-matrix (DM) or Lieb universal density functional

$$F_{\mathsf{DM}}[n] = \min_{\substack{\hat{\Gamma} \in \mathcal{D}_{\mathsf{DM}} \\ \hat{\Gamma}
ightarrow n}} \mathsf{Tr}[(\hat{T} + \hat{W}_{\mathsf{ee}})\hat{\Gamma}]$$

where the minimization is done over admissible *N*-electron ensemble density matrices $\hat{\Gamma} \in \mathcal{D}_{DM} = \{\hat{\Gamma} = \sum_i w_i | \Psi_i \rangle \langle \Psi_i |, 0 \leq w_i \leq 1, \sum_i w_i = 1, \Psi_i \in \mathcal{W}, \langle \Psi_i | \Psi_j \rangle = \delta_{i,j} \}$

- yielding the density n.
 F_{DM}[n] is defined over the same set of N-representable densities D and it is always lower than the Levy-Lieb density functional, i.e. F_{DM}[n] ≤ F[n], with equality in the absence
- of degeneracies.
 Introducing the ground-state energy E₀[v] as a functional of the external potential v,
 Lieb showed that this functional is the Legendre-Fenchel transform of E₀[v]

$$F_{\rm DM}[n] = \sup_{v \in \mathcal{V}} \left(E_0[v] - \int v(\mathbf{r}) n(\mathbf{r}) d\mathbf{r} \right)$$

where the space of potentials is $\mathcal{V}=L^{3/2}(\mathbb{R}^3)+L^\infty(\mathbb{R}^3).$

► F_{DM}[n] has the mathematical advantage of being convex. However, in the remaining, we will only use the Levy-Lieb density functional.

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Kohn-Sham (KS) method: decomposition of the universal functional

▶ In 1965, Kohn and Sham proposed to decompose F[n] as

$$F[n] = T_s[n] + E_{Hxc}[n]$$

 $ightharpoonup T_s[n]$ is the non-interacting kinetic-energy functional:

$$\mathcal{T}_{s}[\textbf{n}] = \min_{\substack{\Phi \in \mathcal{S} \\ \Phi \to \textbf{n}}} \langle \Phi | \, \hat{\mathcal{T}} | \Phi \rangle = \langle \Phi[\textbf{n}] | \, \hat{\mathcal{T}} | \Phi[\textbf{n}] \rangle$$

where the minimization is done over **single-determinant** wave functions

$$\Phi \in \mathcal{S} = \{\Phi = \psi_1 \wedge \psi_2 \wedge \cdots \wedge \psi_N \mid \psi_i \in H^1(\mathbb{R}^3 \times \{\uparrow, \downarrow\}, \mathbb{C}), \langle \psi_i | \psi_j \rangle = \delta_{i,j} \}$$
 yielding the fixed density n . The minimizing single-determinant wave function is called the **KS wave function** and is denoted by $\Phi[n]$.

- ▶ The remaining functional $E_{Hxc}[n]$ is called the **Hartree-exchange-correlation functional**.
- ▶ $T_s[n]$ is still defined over the **entire set of** N-representable **densities** \mathcal{D} because any N-representable density can be obtained from a single-determinant wave function. Therefore, the **Kohn-Sham decomposition does not introduce any approximation**.

Kohn-Sham (KS) method: variational principle

► The **exact ground-state energy** can then be expressed as

$$\begin{split} E_0 &= \min_{n \in \mathcal{D}} \left\{ F[n] + \int v_{\text{ne}}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r} \right\} \\ &= \min_{n \in \mathcal{D}} \left\{ \min_{\substack{\Phi \in \mathcal{S} \\ \Phi \to n}} \langle \Phi | \hat{T} | \Phi \rangle + E_{\text{Hxc}}[n] + \int v_{\text{ne}}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r} \right\} \\ &= \min_{n \in \mathcal{D}} \min_{\substack{\Phi \in \mathcal{S} \\ \Phi \to n}} \left\{ \langle \Phi | \hat{T} + \hat{V}_{\text{ne}} | \Phi \rangle + E_{\text{Hxc}}[n_{\Phi}] \right\} \\ &= \min_{\Phi \in \mathcal{S}} \left\{ \langle \Phi | \hat{T} + \hat{V}_{\text{ne}} | \Phi \rangle + E_{\text{Hxc}}[n_{\Phi}] \right\} \end{split}$$

and the minimizing single-determinant KS wave function gives an exact ground-state density $n_0(\mathbf{r})$.

- Hence, in KS DFT, we replace "min" by "min" which is still a **tremendous** simplification! The advantage of KS DFT over pure DFT is that a major part of the kinetic energy is treated explicitly with the single-determinant wave function Φ .
- KINETIC energy is treated explicitly with the single-determinant wave function

 KS DFT is similar to Hartree-Fock (HF) $E_{\text{HF}} = \min_{n=0}^{\infty} \langle \Phi | \hat{T} + \hat{V}_{\text{ne}} + \hat{W}_{\text{ee}} | \Phi \rangle$

but in KS DFT the exact ground-state energy and density are in principle obtained! 18/95

Kohn-Sham (KS) method: the Hartree-exchange-correlation functional

 $ightharpoonup E_{\mathsf{Hxc}}[n]$ is decomposed as

$$E_{\mathsf{Hxc}}[n] = E_{\mathsf{H}}[n] + E_{\mathsf{xc}}[n]$$

 $E_{H}[n]$ is the **Hartree energy functional**

$$E_{H}[n] = rac{1}{2} \iint rac{n(\mathbf{r}_1)n(\mathbf{r}_2)}{|\mathbf{r}_1 - \mathbf{r}_2|} d\mathbf{r}_1 d\mathbf{r}_2$$

representing the classical electrostatic repulsion energy for the charge distribution $n(\mathbf{r})$ and which is calculated exactly.

and which is calculated exactly.
 E_{xc}[n] is the exchange-correlation energy functional that remains to approximate.
 Assuming Φ[n] is unique (up to a phase factor), this functional is often decomposed as

$$E_{xc}[n] = E_x[n] + E_c[n]$$

where $E_x[n]$ is the exchange energy functional

$$E_{\mathsf{x}}[n] = \langle \Phi[n] | \hat{W}_{\mathsf{ee}} | \Phi[n] \rangle - E_{\mathsf{H}}[n]$$

^[1 ([] ()

and
$$E_c[n]$$
 is the correlation energy functional
$$E_c[n] = \langle \Psi[n] | \hat{T} + \hat{W}_{ee} | \Psi[n] \rangle - \langle \Phi[n] | \hat{T} + \hat{W}_{ee} | \Phi[n] \rangle = T_c[n] + U_c[n]$$

 $E_{c}[n] = \langle \Psi[n] | I + W_{ee} | \Psi[n] \rangle - \langle \Psi[n] | I + W_{ee} | \Psi[n] \rangle = I_{c}[n] + U_{c}[n]$ containing a kinetic contribution $T_{c}[n] = \langle \Psi[n] | \hat{T} | \Psi[n] \rangle - \langle \Phi[n] | \hat{T} | \Phi[n] \rangle$

containing a kinetic contribution $I_c[n] = \langle \Psi[n] | I | \Psi[n] \rangle - \langle \Phi[n] | I | \Phi[n] \rangle$ and a potential contribution $U_c[n] = \langle \Psi[n] | \hat{W}_{ee} | \Psi[n] \rangle - \langle \Phi[n] | \hat{W}_{ee} | \Phi[n] \rangle$.

The Kohn-Sham equations (1/2) The single determinant Φ is constructed from a set of N orthonormal occupied

spin-orbitals $\psi_i(\mathbf{x}) = \varphi_i(\mathbf{r})\delta_{\sigma_i,\sigma}$.

► The **total energy** to be minimized is

$$E[\{\varphi_i\}] = \sum_{i=1}^{N} \int \varphi_i^*(\mathbf{r}) \left(-\frac{1}{2} \nabla^2 + v_{\text{ne}}(\mathbf{r}) \right) \varphi_i(\mathbf{r}) d\mathbf{r} + E_{\text{Hxc}}[n]$$

and the density is

$$n(\mathbf{r}) = \sum_{i=1}^{N} |\varphi_i(\mathbf{r})|^2$$

For minimizing over the orbitals $\{\varphi_i\}$ with the constraint of keeping the orbitals orthonormalized, we introduce the **Lagrangian**

$$\mathcal{L}[\{\varphi_i\}] = E[\{\varphi_i\}] - \sum_{i=1}^{N} \varepsilon_i \left(\int \varphi_i^*(\mathbf{r}) \varphi_i(\mathbf{r}) d\mathbf{r} - 1 \right)$$

where ε_i is the Lagrange multiplier associated with the normalization condition of $\varphi_i(\mathbf{r})$.

The Lagrangian must be **stationary** with respect to variations of the orbitals $\varphi_i(\mathbf{r})$

$$\frac{\delta \mathcal{L}}{\delta \varphi_i^*(\mathbf{r})} = 0$$

Interlude: Review on functional derivatives

► For a **functional** $F: f \mapsto F[f]$ of the function $f: x \mapsto f(x)$, an infinitesimal variation δf of f leads to an infinitesimal variation of F which can be expressed as

$$\delta F[f] = \int \frac{\delta F[f]}{\delta f(x)} \delta f(x) dx$$

This defines the **functional derivative** of F[f] with respect to f(x): $\frac{\delta F[f]}{\delta f(x)}$

Remark: For a function $F(f_1, f_2, ...)$ of several variables $f_1, f_2, ...,$ we have

 $dF = \sum_i rac{\partial F}{\partial f_i} \ df_i$

 $\delta F[f]/\delta f(x)$ is the analog of $\partial F/\partial f_i$ for the case of an infinite continuous number of variables.

For a functional F[f] of a function f[g](x) which is itself a functional of another function g(x), we have the **chain rule**

$$\frac{\delta F}{\delta g(x)} = \int \frac{\delta F}{\delta f(x')} \frac{\delta f(x')}{\delta g(x)} \mathrm{d}x'$$

Remark: It is the analog of the chain rule for a function $F(f_1, f_2, ...)$ of several variables $f_j(g_1, g_2, ...)$ which are themselves functions of other variables $g_1, g_2, ...$

$$\frac{\partial F}{\partial g_i} = \sum_j \frac{\partial F}{\partial f_j} \frac{\partial f_j}{\partial g_i}$$

The Kohn-Sham equations (2/2)

▶ We find for the functional derivative of the Lagrangian

$$0 = \frac{\delta \mathcal{L}}{\delta \varphi_i^*(\mathbf{r})} = \left(-\frac{1}{2}\nabla^2 + \nu_{\text{ne}}(\mathbf{r})\right)\varphi_i(\mathbf{r}) + \frac{\delta E_{\text{Hxc}}[\mathbf{n}]}{\delta \varphi_i^*(\mathbf{r})} - \varepsilon_i \varphi_i(\mathbf{r})$$

▶ We calculate the term $\delta E_{\text{Hxc}}[n]/\delta \varphi_i^*(\mathbf{r})$ using the chain rule

$$\frac{\delta E_{\mathsf{Hxc}}[n]}{\delta \varphi_i^*(\mathbf{r})} = \int \frac{\delta E_{\mathsf{Hxc}}[n]}{\delta n(\mathbf{r}')} \frac{\delta n(\mathbf{r}')}{\delta \varphi_i^*(\mathbf{r})} \mathsf{d}\mathbf{r}' = \mathsf{v}_{\mathsf{Hxc}}(\mathbf{r}) \varphi_i(\mathbf{r})$$

where we have used $\delta n(\mathbf{r}')/\delta \varphi_i^*(\mathbf{r}) = \varphi_i(\mathbf{r})\delta(\mathbf{r} - \mathbf{r}')$ and we have introduced the Hartree-exchange-correlation potential $v_{\text{Hxc}}(\mathbf{r})$

$$v_{\mathsf{Hxc}}(\mathbf{r}) = \frac{\delta E_{\mathsf{Hxc}}[n]}{\delta n(\mathbf{r})}$$

which is itself a functional of the density.

We arrive at the KS equations

$$\left(-\frac{1}{2}\nabla^2 + v_{\mathsf{ne}}(\mathbf{r}) + v_{\mathsf{Hxc}}(\mathbf{r})\right)\varphi_i(\mathbf{r}) = \varepsilon_i\varphi_i(\mathbf{r})$$

The orbitals $\varphi_i(\mathbf{r})$ are called the KS orbitals and ε_i are the KS orbital energies.

The Kohn-Sham equations and the Kohn-Sham Hamiltonian

▶ The KS orbitals are eigenfunctions of the KS one-electron Hamiltonian

$$h_{\mathsf{s}} = -rac{1}{2}
abla^2 +
u_{\mathsf{s}}(\mathbf{r})$$

where $v_s(\mathbf{r}) = v_{ne}(\mathbf{r}) + v_{Hxc}(\mathbf{r})$ is the KS potential.

- Mathematically, the KS equations are a set of coupled self-consistent equations since the potential $v_{Hxc}(\mathbf{r})$ depends on all the occupied orbitals $\{\varphi_i\}_{i=1,...N}$ through the density.
- Physically, h_s defines the KS system which is a system of N non-interacting electrons in an effective external potential $v_s(\mathbf{r})$ ensuring that its ground-state density $n(\mathbf{r})$ is the same as the exact ground-state density $n_0(\mathbf{r})$ of the physical system of N interacting electrons.
- ▶ The KS equations also defines virtual KS orbitals $\{\varphi_a\}_{a>N+1}$.
- $h_s(\mathbf{r})$ is a self-adjoint operator and thus the set of KS orbitals can be chosen as orthonormal.

The Hartree-exchange-correlation potential

▶ To define $v_{\text{Hxc}}(\mathbf{r}) = \delta E_{\text{Hxc}}[n]/\delta n(\mathbf{r})$, we have assumed a form of differentiability of $E_{\text{Hxc}}[n]$. This can in fact only hold on a restricted set of densities. This is known as the v-representability problem.

► The KS potential is defined only up to an additive constant. For atomic and molecular ground-state densities, we choose the constant so that the potential vanishes at infinity: $\lim_{|\mathbf{r}| \to \infty} v_s(\mathbf{r}) = 0$

Following the decomposition of $E_{\rm Hxc}[n]$, the potential $v_{\rm Hxc}({\bf r})$ is also decomposed as $v_{\rm Hxc}({\bf r}) = v_{\rm H}({\bf r}) + v_{\rm xc}({\bf r})$

with the Hartree potential $v_{\rm H}({\bf r}) = \frac{\delta E_{\rm H}[n]}{\delta n({\bf r})} = \int \frac{n({\bf r}')}{|{\bf r}-{\bf r}'|} {\rm d}{\bf r}'$ and the exchange-correlation potential $v_{\rm xc}({\bf r}) = \delta E_{\rm xc}[n]/\delta n({\bf r})$

The potential
$$v_{xc}(\mathbf{r})$$
 can be decomposed as $v_{xc}(\mathbf{r}) = v_x(\mathbf{r}) + v_c(\mathbf{r})$ with the **exchange potential** $v_x(\mathbf{r}) = \delta E_x[n]/\delta n(\mathbf{r})$

and the correlation potential $v_c(\mathbf{r}) = \delta E_c[n]/\delta n(\mathbf{r})$

Practical calculations in an atomic basis (1/3)

▶ We consider a **basis of** M **atom-centered functions** $\{\chi_{\nu}\}$, e.g. GTO basis functions. The orbitals are expanded as

$$arphi_i(\mathbf{r}) = \sum_{
u=1}^M c_{
u i} \; \chi_
u(\mathbf{r})$$

Inserting this expansion in the KS equations

$$h_{s}\varphi_{i}(\mathbf{r})=\varepsilon_{i}\varphi_{i}(\mathbf{r})$$

and multiplying on the left by $\chi_{\mu}^*(\mathbf{r})$ and integrating over \mathbf{r} , we arrive at the familiar SCF generalized eigenvalue equation

$$\sum_{\nu=1}^{M} F_{\mu\nu} c_{\nu i} = \varepsilon_i \sum_{\nu=1}^{M} S_{\mu\nu} c_{\nu i}$$

where $F_{\mu\nu}=\int\chi_{\mu}^*(\mathbf{r})h_s\chi_{\nu}(\mathbf{r})d\mathbf{r}$ are the elements of the KS Fock matrix and $S_{\mu\nu}=\int\chi_{\mu}^*(\mathbf{r})\chi_{\nu}(\mathbf{r})d\mathbf{r}$ are the elements of the overlap matrix.

Practical calculations in an atomic basis (2/3)

▶ The Fock matrix is calculated as $F_{\mu\nu} = h_{\mu\nu} + J_{\mu\nu} + V_{\text{xc.}\,\mu\nu}$

$$m{h}_{\mu
u}$$
 are the one-electron integrals: $h_{\mu
u}=\int\chi_{\mu}^*({f r})\left(-rac{1}{2}
abla^2+
u_{
m ne}({f r})
ight)\chi_{
u}({f r}){
m d}{f r}$

 $J_{\mu\nu}$ is the Hartree potential matrix:

$$J_{\mu
u} = \int \chi_{\mu}^*(\mathbf{r}) \mathsf{v}_\mathsf{H}(\mathbf{r}) \chi_{
u}(\mathbf{r}) \mathsf{d}\mathbf{r} = \sum_{\lambda=1}^M \sum_{\gamma=1}^M P_{\gamma\lambda}(\chi_{\mu}\chi_{
u}|\chi_{\lambda}\chi_{\gamma})$$

where $(\chi_{\mu}\chi_{\nu}|\chi_{\lambda}\chi_{\gamma}) = \iint \frac{\chi_{\mu}^{*}(\mathbf{r}_{1})\chi_{\nu}(\mathbf{r}_{1})\chi_{\lambda}^{*}(\mathbf{r}_{2})\chi_{\gamma}(\mathbf{r}_{2})}{|\mathbf{r}_{1}-\mathbf{r}_{2}|_{\mu}} d\mathbf{r}_{1}d\mathbf{r}_{2}$ are the two-electron integrals (in chemists' notation) and $P_{\gamma\lambda}=\sum c_{\gamma i}c_{\lambda i}^*$ is the density matrix.

$$m{V}_{{
m xc},\mu
u}$$
 is the exchange-correlation potential matrix: $m{V}_{{
m xc},\mu
u}=\int\chi_{\mu}^*({f r})m{v}_{{
m xc}}({f r})\chi_{
u}({f r}){
m d}{f r}$

The total electronic energy is calculated as

$$E = \sum_{i=1}^{M} \sum_{j=1}^{M} P_{\nu\mu} h_{\mu\nu} + \frac{1}{2} \sum_{j=1}^{M} \sum_{j=1}^{M} P_{\nu\mu} J_{\mu\nu} + E_{xc}$$

► The density is calculated as $n(\mathbf{r}) = \sum_{m}^{m} \sum_{m}^{m} P_{\gamma\lambda} \chi_{\gamma}(\mathbf{r}) \chi_{\lambda}^{*}(\mathbf{r})$

Practical calculations in an atomic basis (3/3)

In the simplest approximation, the exchange-correlation energy functional has a local form

$$E_{xc}^{local} = \int f(n(\mathbf{r})) d\mathbf{r}$$

where $f(n(\mathbf{r}))$ has a complicated nonlinear dependence on the density $n(\mathbf{r})$.

▶ For example, in the local-density approximation (LDA), the exchange energy is

$$E_{x}^{LDA} = C_{x} \int n(\mathbf{r})^{4/3} d\mathbf{r}$$

where C_x is a constant, and the exchange potential is

$$v_{x}^{LDA}(\mathbf{r}) = \frac{4}{3}C_{x}n(\mathbf{r})^{1/3}$$

► Therefore, the integrals cannot be calculated analytically, but are instead evaluated by numerical integration on a grid

$$V_{ ext{xc},\mu
u}pprox \sum_{k} w_k \; \chi_{\mu}^*(\mathbf{r}_k) v_{ ext{xc}}(\mathbf{r}_k) \chi_{
u}(\mathbf{r}_k) \qquad ext{and} \qquad E_{ ext{xc}}^{ ext{local}}pprox \sum_{k} w_k \; f(n(\mathbf{r}_k))$$

where \mathbf{r}_k and w_k are quadrature points and weights. For molecules, the multicenter numerical integration scheme of Becke (1988) is generally used.

Extension to spin density-functional theory (1/2)

For dealing with an external magnetic field, DFT has been extended from the total density to **spin-resolved densities** (von Barth and Hedin, 1972)

$$n_{\sigma}(\mathbf{r}) = N \int \cdots \int |\Psi(\mathbf{r}\sigma, \mathbf{x}_{2}, ..., \mathbf{x}_{N})|^{2} d\mathbf{x}_{2}...d\mathbf{x}_{N}$$
 with $\sigma \in \{\uparrow, \downarrow\}$

- which integrate to the numbers of σ -spin electrons, i.e. $\int n_{\sigma}(\mathbf{r})d\mathbf{r} = N_{\sigma}$.
- Without magnetic fields, this is in principle not necessary. In practice, the dependence on the spin densities allows one to construct more accurate approximate exchange-correlation functionals for open-shell systems.

where the search is over wave functions $\Psi \in \mathcal{W}$ with $N = N_{\uparrow} + N_{\downarrow}$ electrons and which

► The universal density functional is now defined as

$$F[n_{\uparrow},n_{\downarrow}] = \min_{egin{subarray}{c} \Psi \in \mathcal{W} \ \psi
ightarrow n_{\uparrow},n_{\downarrow} \ \end{array}} \langle \Psi | \hat{\mathcal{T}} + \hat{\mathcal{W}}_{\mathsf{ee}} | \Psi
angle$$

- yield fixed spin densities.
- A **KS method** is obtained by decomposing $F[\mathit{n}_{\uparrow},\mathit{n}_{\downarrow}]$ as

 $F[n_{\uparrow},n_{\downarrow}] = T_{\rm s}[n_{\uparrow},n_{\downarrow}] + E_{\rm H}[n] + E_{\rm xc}[n_{\uparrow},n_{\downarrow}]$ where $T_{\rm s}[n_{\uparrow},n_{\downarrow}]$ is defined with a constrained search over (spin-unrestricted) Slater

determinants
$$\Phi$$

$$T_{\mathsf{s}}[n_{\uparrow},n_{\downarrow}] = \min_{\substack{\Phi \in \mathcal{S} \\ \Phi \neq \mathcal{S} \\ n_{\downarrow}}} \langle \Phi | \hat{T} | \Phi \rangle$$

The second area of a second as

Extension to spin density-functional theory (2/2)

► The **exact ground-state energy** is expressed as

$$E_0 = \min_{\Phi \in \mathcal{S}} \left\{ \langle \Phi | \hat{T} + \hat{V}_{\mathsf{ne}} | \Phi \rangle + E_{\mathsf{H}}[n_{\Phi}] + E_{\mathsf{xc}}[n_{\uparrow,\Phi}, n_{\downarrow,\Phi}] \right\}$$

Writing the spatial orbitals of the determinant as $\varphi_{i\sigma}(\mathbf{r})$ (with indices explicitly including spin now), we have now the **spin-dependent KS equations**

$$\left(-rac{1}{2}
abla^2 + v_{\mathsf{ne}}(\mathbf{r}) + v_{\mathsf{H}}(\mathbf{r}) + v_{\mathsf{xc},\sigma}(\mathbf{r})
ight)arphi_{i\sigma}(\mathbf{r}) = arepsilon_{i\sigma}arphi_{i\sigma}(\mathbf{r})$$

with the spin-dependent exchange-correlation potential and density

$$v_{ ext{xc},\sigma}(\mathbf{r}) = rac{\delta E_{ ext{xc}}[n_{\uparrow},n_{\downarrow}]}{\delta n_{\sigma}(\mathbf{r})} ext{ and } n_{\sigma}(\mathbf{r}) = \sum_{i=1}^{N_{\sigma}} |arphi_{i\sigma}(\mathbf{r})|^2$$

▶ The spin-dependent exchange functional $E_x[n_\uparrow, n_\downarrow]$ can be obtained from the spin-independent exchange functional $E_x[n]$ with the **spin-scaling relation**

ange functional
$$E_{\times}[n]$$
 with the **spin-scaling relation**
$$E_{\times}[n_{\uparrow},n_{\downarrow}]=\frac{1}{2}\left(E_{\times}[2n_{\uparrow}]+E_{\times}[2n_{\downarrow}]\right)$$

Therefore, any approximation for the spin-independent exchange functional $E_x[n]$ can be easily extended to an approximation for the spin-dependent exchange functional

 $E_{\mathsf{x}}[n_{\uparrow}, n_{\downarrow}]$. Unfortunately, there is no such relation for the correlation functional.

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 The Kohn-Sham equations
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 - Generalized Kohn-Sham method

Generalized Kohn-Sham method

An important extension of the KS method is the **generalized Kohn-Sham** (GKS) method (1996) in which the universal density functional F[n] is decomposed as

$$F[n] = \min_{\substack{\Phi \in \mathcal{S} \\ \Phi \to n}} \left\{ \langle \Phi | \hat{T} | \Phi \rangle + E_{\mathsf{H}}[n_{\Phi}] + S[\Phi] \right\} + \bar{S}[n]$$
 where $S[\Phi]$ is any (reasonable) functional of a single-determinant wave function $\Phi \in \mathcal{S}$

and $\bar{S}[n]$ is the complementary density functional. E.g., in hybrids, $S[\Phi] = aE_x^{HF}[\Phi]$.

Defining the **GKS exchange-correlation functional**, $E_{xc}^S[\Phi] = S[\Phi] + \bar{S}[n_{\Phi}]$, we can

express the **exact ground-state energy** as
$$E_0 = \min_{\Phi \in \mathcal{S}} \left\{ \langle \Phi | \hat{T} + \hat{V}_{\text{ne}} | \Phi \rangle + E_{\text{H}}[\textit{n}_{\Phi}] + \textit{E}_{\text{xc}}^{\mathcal{S}}[\Phi] \right\}$$

and any minimizing single-determinant wave function gives a ground-state density $n_0(\mathbf{r})$.

 $\left(-\frac{1}{2}\nabla^2 + v_{\rm ne}(\mathbf{r}) + v_{\rm H}(\mathbf{r}) + v_{\bar{5}}(\mathbf{r})\right)\varphi_{i\sigma}(\mathbf{r}) + \frac{\delta S[\Phi]}{\delta\varphi_{i\sigma}^*(\mathbf{r})} = \varepsilon_{i\sigma}\varphi_{i\sigma}(\mathbf{r})$ where $v_{\bar{5}}(\mathbf{r}) = \delta\bar{S}[n]/\delta n(\mathbf{r})$ is a local potential and $\delta S[\Phi]/\delta\varphi_{i\sigma}^*(\mathbf{r})$ generates a

where $v_{\bar{s}}(\mathbf{r}) = \delta \bar{S}[n]/\delta n(\mathbf{r})$ is a local potential and $\delta S[\Phi]/\delta \varphi_{i\sigma}^*(\mathbf{r})$ generates a one-electron (possibly nonlocal) operator.

The GKS method gives **much more freedom** than the KS method (which corresponds to the special case $S[\Phi] = 0$).

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- Usual approximations for the exchange-correlation energy
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The exchange-correlation hole

▶ The **pair** density associated with the wave function $\Psi[n]$ is

$$n_2(\mathbf{r}_1,\mathbf{r}_2) = N(N-1) \int \cdots \int |\Psi[n](\mathbf{x}_1,\mathbf{x}_2,...,\mathbf{x}_N)|^2 d\sigma_1 d\sigma_2 d\mathbf{x}_3...d\mathbf{x}_N$$

which is a functional of the density. It is normalized to the number of electron pairs: $\iint n_2(\mathbf{r}_1, \mathbf{r}_2) d\mathbf{r}_1 d\mathbf{r}_2 = N(N-1).$ It is proportional to the probability density of finding two electrons at positions $(\mathbf{r}_1, \mathbf{r}_2)$ with all the other electrons anywhere.

 $\langle \Psi[\emph{n}] | \hat{\mathcal{W}}_{\mathsf{ee}} | \Psi[\emph{n}]
angle = rac{1}{2} \iint rac{\emph{n}_2(extsf{r}_1, extsf{r}_2)}{| extsf{r}_1 - extsf{r}_2|} \mathsf{d} extsf{r}_1 \mathsf{d} extsf{r}_2$

It can be used to express the electron-electron interaction energy

Mirroring the decomposition of
$$E_{Hxc}[n]$$
, the pair density can be decomposed as
$$n_2(\mathbf{r}_1,\mathbf{r}_2) = n(\mathbf{r}_1)n(\mathbf{r}_2) + n_{2,xc}(\mathbf{r}_1,\mathbf{r}_2)$$

electrons correlation effects

We also introduce the **exchange-correlation hole** $n_{xc}(\mathbf{r}_1, \mathbf{r}_2)$ by

$$n_{2,xc}(\mathbf{r}_1,\mathbf{r}_2) = n(\mathbf{r}_1)n_{xc}(\mathbf{r}_1,\mathbf{r}_2)$$

It can be interpreted as the modification due to exchange and correlation effects of the conditional probability of finding an electron at \mathbf{r}_2 knowing that one has been found at \mathbf{r}_1 .

We have the exact constraints: $n_{xc}(\mathbf{r}_1,\mathbf{r}_2) \geq -n(\mathbf{r}_2)$ and $\int n_{xc}(\mathbf{r}_1,\mathbf{r}_2)d\mathbf{r}_2 = -1$

The exchange hole

 \blacktriangleright Similarly, we define the **KS pair density** associated with the KS single determinant $\Phi[n]$

$$n_{2,KS}(\mathbf{r}_1,\mathbf{r}_2) = N(N-1) \int \cdots \int |\Phi[n](\mathbf{x}_1,\mathbf{x}_2,...,\mathbf{x}_N)|^2 d\sigma_1 d\sigma_2 d\mathbf{x}_3...d\mathbf{x}_N$$

It can be decomposed as

$$n_{2,KS}(\mathbf{r}_1,\mathbf{r}_2) = n(\mathbf{r}_1)n(\mathbf{r}_2) + n_{2,x}(\mathbf{r}_1,\mathbf{r}_2)$$

and we introduce the **exchange hole** $n_x(\mathbf{r}_1, \mathbf{r}_2)$ by $n_{2,x}(\mathbf{r}_1, \mathbf{r}_2) = n(\mathbf{r}_1)n_x(\mathbf{r}_1, \mathbf{r}_2)$

which satisfies the exact constraints: $n_x(\mathbf{r}_1, \mathbf{r}_2) \ge -n(\mathbf{r}_2)$ and $\int n_x(\mathbf{r}_1, \mathbf{r}_2) d\mathbf{r}_2 = -1$ and $n_x(\mathbf{r}_1, \mathbf{r}_2) \le 0$

The exchange energy functional is the electrostatic interaction energy between an electron and its exchange hole:

$$E_{\mathbf{x}}[n] = \frac{1}{2} \iint \frac{n(\mathbf{r}_1)n_{\mathbf{x}}(\mathbf{r}_1, \mathbf{r}_2)}{|\mathbf{r}_1 - \mathbf{r}_2|} d\mathbf{r}_1 d\mathbf{r}_2 = \int n(\mathbf{r}_1)\varepsilon_{\mathbf{x}}[n](\mathbf{r}_1) d\mathbf{r}_1$$

where $\varepsilon_{\mathbf{x}}[n](\mathbf{r}_1)$ is the exchange energy per particle. In approximate exchange density functionals, the quantity $\varepsilon_{\mathbf{x}}[n](\mathbf{r}_1)$ is usually what is approximated.

The correlation hole

The correlation hole is defined as the difference

$$n_{c}(\mathbf{r}_{1},\mathbf{r}_{2}) = n_{xc}(\mathbf{r}_{1},\mathbf{r}_{2}) - n_{x}(\mathbf{r}_{1},\mathbf{r}_{2})$$

and satisfies the sum rule

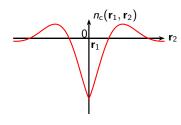
$$\int n_{\rm c}(\mathbf{r}_1,\mathbf{r}_2)\mathrm{d}\mathbf{r}_2=0$$

which implies that the correlation hole has negative and positive contributions.

The potential contribution to the correlation energy can be written in terms of the correlation hole

$$U_{c}[n] = \frac{1}{2} \iint \frac{n(\mathbf{r}_{1})n_{c}(\mathbf{r}_{1}, \mathbf{r}_{2})}{|\mathbf{r}_{1} - \mathbf{r}_{2}|} d\mathbf{r}_{1} d\mathbf{r}_{2}$$

But in order to express the total correlation energy $E_c[n] = T_c[n] + U_c[n]$ in a similar form, we need to introduce the adiabatic-connection formalism.



The adiabatic connection (1/3)

- The idea of the adiabatic connection is to have a continuous path between the non-interacting KS system and the physical system while keeping the ground-state density constant.
- ightharpoonup For this, we introduce a Hamiltonian depending on a **coupling constant** λ which switches on the electron-electron interaction

$$\hat{H}^{\lambda} = \hat{T} + \lambda \hat{W}_{ee} + \hat{V}^{\lambda}$$

where \hat{V}^{λ} is the external local potential imposing that the ground-state density is the same as the ground-state density of the physical system for all λ , i.e. $n^{\lambda}(\mathbf{r}) = n_0(\mathbf{r}), \forall \lambda$.

▶ By varying λ , we connect the KS non-interacting system ($\lambda = 0$) to the physical interacting system ($\lambda = 1$):

$$\underbrace{\hat{\mathcal{H}}^{\lambda=0}}_{\text{KS non-interacting}} \xleftarrow{0 \leq \lambda \leq 1} \underbrace{\hat{\mathcal{H}}^{\lambda=1}}_{\text{Physical interacting}}$$

 \blacktriangleright We define a universal functional for each value of the parameter λ

$$F^{\lambda}[n] = \min_{\substack{\Psi \in \mathcal{W} \\ \Psi \to n}} \langle \Psi | \, \hat{T} + \lambda \, \hat{W}_{\text{ee}} | \Psi \rangle = \langle \Psi^{\lambda}[n] | \, \hat{T} + \lambda \, \hat{W}_{\text{ee}} | \Psi^{\lambda}[n] \rangle$$

The adiabatic connection (2/3)

▶ The functional $F^{\lambda}[n]$ can be decomposed as

$$F^{\lambda}[n] = T_{s}[n] + E_{H}^{\lambda}[n] + E_{x}^{\lambda}[n] + E_{c}^{\lambda}[n]$$

▶ $E_{\rm H}^{\lambda}[n]$ and $E_{\rm x}^{\lambda}[n]$ are the Hartree and exchange functionals associated with the interaction $\lambda \hat{W}_{\rm ee}$ and are simply linear in λ

$$E_{\mathrm{H}}^{\lambda}[n] = \lambda E_{\mathrm{H}}[n]$$
 and $E_{\mathrm{x}}^{\lambda}[n] = \lambda E_{\mathrm{x}}[n]$

▶ The correlation functional $E_c^{\lambda}[n]$ is nonlinear in λ

$$\textit{E}_{c}^{\lambda}[\textit{n}] = \langle \Psi^{\lambda}[\textit{n}] | \, \hat{\textit{T}} + \lambda \hat{\textit{W}}_{ee} | \Psi^{\lambda}[\textit{n}] \rangle - \langle \Phi[\textit{n}] | \, \hat{\textit{T}} + \lambda \hat{\textit{W}}_{ee} | \Phi[\textit{n}] \rangle$$

▶ We can get rid of \hat{T} by taking the derivative with respect to λ and using the Hellmann-Feynman theorem for the wave function $\Psi^{\lambda}[n]$

$$\frac{\partial E_{\rm c}^{\lambda}[n]}{\partial \lambda} = \langle \Psi^{\lambda}[n] | \hat{W}_{\rm ee} | \Psi^{\lambda}[n] \rangle - \langle \Phi[n] | \hat{W}_{\rm ee} | \Phi[n] \rangle$$

The adiabatic connection (3/3)

▶ Reintegrating over λ from 0 to 1, and using $E_c^{\lambda=1}[n] = E_c[n]$ and $E_c^{\lambda=0}[n] = 0$ (assuming no degeneracies at $\lambda = 0$), we arrive at the **adiabatic-connection formula**

$$E_{\mathsf{c}}[n] = \int_0^1 \mathsf{d}\lambda \; \langle \Psi^\lambda[n] | \hat{\mathcal{W}}_{\mathsf{ee}} | \Psi^\lambda[n]
angle - \langle \Phi[n] | \hat{\mathcal{W}}_{\mathsf{ee}} | \Phi[n]
angle$$

▶ Introducing the correlation hole $n_c^{\lambda}(\mathbf{r}_1, \mathbf{r}_2)$ associated with the wave function $\Psi^{\lambda}[n]$, the adiabatic-connection formula can also be written as

$$E_{c}[n] = \frac{1}{2} \int_{0}^{1} d\lambda \iint \frac{n(\mathbf{r}_{1}) n_{c}^{\lambda}(\mathbf{r}_{1}, \mathbf{r}_{2})}{|\mathbf{r}_{1} - \mathbf{r}_{2}|} d\mathbf{r}_{1} d\mathbf{r}_{2}$$

▶ Introducing the λ -integrated correlation hole $\bar{n}_c(\mathbf{r}_1,\mathbf{r}_2) = \int_0^1 d\lambda \ n_c^{\lambda}(\mathbf{r}_1,\mathbf{r}_2)$, we finally write

$$E_{c}[n] = \frac{1}{2} \iint \frac{n(\mathbf{r}_1)\overline{n}_{c}(\mathbf{r}_1, \mathbf{r}_2)}{|\mathbf{r}_1 - \mathbf{r}_2|} d\mathbf{r}_1 d\mathbf{r}_2 = \int n(\mathbf{r}_1)\varepsilon_{c}[n](\mathbf{r}_1) d\mathbf{r}_1$$

where $\varepsilon_{\rm c}[n]({\bf r}_1)$ is the correlation energy per particle, which is the quantity usually approximated in practice.

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Uniform coordinate scaling (1/2)

▶ We consider a **norm-preserving uniform coordinate scaling** in the *N*-electron wave function along the adiabatic connection $Ψ^{\lambda}[n]$ (ignoring untouched spin variables)

$$\Psi_{\gamma}^{\lambda}[n](\mathbf{r}_{1},...,\mathbf{r}_{N}) = \frac{\gamma^{3N/2}}{\gamma^{N/2}}\Psi^{\lambda}[n](\gamma\mathbf{r}_{1},...,\gamma\mathbf{r}_{N})$$

where $\gamma > 0$ is a scaling factor.

► The scaled wave function $\Psi^{\lambda}_{\gamma}[n]$ yields the scaled density

$$n_{\gamma}(\mathbf{r}) = \gamma^{3} n(\gamma \mathbf{r})$$
 (with $\int n_{\gamma}(\mathbf{r}) d\mathbf{r} = \int n(\mathbf{r}) d\mathbf{r} = N$)

and minimizes $\langle \Psi | \hat{T} + \lambda \gamma \hat{W}_{ee} | \Psi \rangle$ since it can be shown that

$$\langle \Psi_{\gamma}^{\lambda}[n]|\hat{T} + \lambda_{\gamma}\hat{W}_{ee}|\Psi_{\gamma}^{\lambda}[n]\rangle = \frac{\gamma^{2}}{2}\langle \Psi^{\lambda}[n]|\hat{T} + \lambda\hat{W}_{ee}|\Psi^{\lambda}[n]\rangle$$

We thus conclude

$$\Psi^{\lambda}_{\gamma}[n] = \Psi^{\lambda\gamma}[n_{\gamma}]$$
 or, equivalently, $\Psi^{\lambda/\gamma}_{\gamma}[n] = \Psi^{\lambda}[n_{\gamma}]$

and for the universal density functional

$$F^{\lambda\gamma}[n_{\gamma}] = \gamma^2 F^{\lambda}[n]$$
 or, equivalently, $F^{\lambda}[n_{\gamma}] = \gamma^2 F^{\lambda/\gamma}[n]$

Uniform coordinate scaling (2/2)

At $\lambda = 0$, we find the scaling relation of the KS single-determinant wave function

$$\Phi[n_{\gamma}] = \Phi_{\gamma}[n]$$

▶ This directly leads to the scaling relations for $T_s[n]$, $E_H[n]$, and $E_x[n]$

$$T_{\rm s}[n_{\gamma}] = \gamma^2 T_{\rm s}[n]$$
 and $E_{\rm H}[n_{\gamma}] = \gamma E_{\rm H}[n]$ and $E_{\rm x}[n_{\gamma}] = \gamma E_{\rm x}[n]$

▶ However, $E_c[n]$ has the more complicated scaling (as F[n])

$$E_{\rm c}^{\lambda}[n_{\gamma}] = \gamma^2 E_{\rm c}^{\lambda/\gamma}[n]$$

and, in particular for $\lambda = 1$,

$$E_{\rm c}[n_{\gamma}] = {\gamma^2 E_{\rm c}^{1/\gamma}[n]}$$

High- and low-density limits

▶ In the **high-density limit** $(\gamma \to \infty)$, the correlation functional goes to a constant, for nondegenerate KS systems,

$$\lim_{\gamma\to\infty}E_{\rm c}[n_{\gamma}]=E_{\rm c}^{\rm GL2}[n]$$

where $E_{\rm c}^{\rm GL2}[n]$ is the second-order Görling-Levy (GL2) correlation energy.

- ▶ This is also called the **weak-correlation limit** since $E_c[n] \ll E_x[n]$.
- Atomic and molecular systems are often close to the high-density limit. E.g., for the ground-state density of He, $E_c[n] = -0.0421$ a.u. and $\lim_{\gamma \to \infty} E_c[n_{\gamma}] = -0.0467$ a.u..
- In the **low-density limit** ($\gamma \to 0$), the Hartree-exchange-correlation functional goes to zero linearly in γ

$$E_{\mathsf{Hxc}}[n_{\gamma}] \underset{\gamma \to 0}{\sim} \gamma \ W_{\mathsf{ee}}^{\mathsf{SCE}}[n]$$

where $W_{\text{ee}}^{\text{SCE}}[n] = \inf_{\substack{\Psi \in \mathcal{W} \\ \Psi \mid \mathcal{W}}} \langle \Psi | \hat{W}_{\text{ee}} | \Psi \rangle$ is the strictly-correlated-electron (SCE) functional.

- ▶ This limit corresponds to a Wigner crystallization.
- ▶ This is also called the **strong-correlation limit** because $E_c[n] \sim E_x[n]$.
- ▶ Calculation of $W_{\text{ee}}^{\text{SCE}}[n]$ is computationally involved but has been done for a few systems (Seidl, Gori-Giorgi, ...).

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One-orbital spatial regions and self-interaction

▶ For one-electron densities $n_{1e}(\mathbf{r}) = |\varphi_1(\mathbf{r})|^2$ where φ_1 is the unique occupied KS orbital, we have

$$E_{x}[n_{1e}] = -E_{H}[n_{1e}]$$
 and $E_{c}[n_{1e}] = 0$

► For opposite-spin two-electron densities $n_{2\mathrm{e}}^{\uparrow\downarrow}(\mathbf{r}) = 2|\varphi_1(\mathbf{r})|^2$ where φ_1 is the unique doubly occupied KS orbital, we have

$$E_{ extsf{x}}[n_{2 extsf{e}}^{\uparrow\downarrow}] = -rac{1}{2}E_{ extsf{H}}[n_{2 extsf{e}}^{\uparrow\downarrow}]$$

- For systems with more electrons, similar relations apply locally in one-orbital spatial regions, i.e. in regions where only one occupied KS orbital is not zero. This situation can be approximately realized in chemical systems (unpair electron in a radical, and electron pair in a single covalent bond, in a lone pair, or in a core orbital).
- ▶ If approximate exchange and correlation density functionals do not satisfy these constraints, we say that they introduce a **self-interaction error**.

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Lieb-Oxford lower bound

Lieb and Oxford derived a useful lower bound which can be expressed as

$$E_{\mathrm{x}}[n] \geq E_{\mathrm{xc}}[n] \geq -C_{\mathrm{LO}} \int \!\! n(\mathbf{r})^{4/3} \mathrm{d}\mathbf{r}$$

where the optimal (i.e., smallest) constant C_{LO} (independent of the electron number N) has been narrowed to $1.4442 \le C_{LO} \le 1.5765$.

- ▶ This bound is approached in the low-density limit.
- ► For one-electron densities and opposite-spin two-electron densities, specific tigher bounds (i.e., with smaller C_{LO}) are known.

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The HOMO energy and the ionization energy

- ► For clarity, we will explicitly indicate the dependence on the electron number *N* in this section.
- For finite systems, the exact ground-state density of a N-electron system decays exponentially for $r=|{\bf r}|\to\infty$ with an exponent related to the **ionization energy** $I_N=E_0^{N-1}-E_0^N$

$$n^N(\mathbf{r}) \underset{r \to \infty}{\propto} e^{-2\sqrt{2I_N} r}$$

Choosing the constant in the KS potential so that it goes to zero at infinity, i.e. $\lim_{|\mathbf{r}|\to\infty} v_s^N(\mathbf{r}) = 0$, it can be shown the density calculated from the KS orbitals decays exponentially with an exponent related to the **HOMO energy** ε_H^N

$$n^{N}(\mathbf{r}) = \sum_{i=1}^{N} |\varphi_{i}(\mathbf{r})|^{2} \underset{r \to \infty}{\propto} e^{-2\sqrt{-2\varepsilon_{\mathrm{H}}^{N}} r}$$

► This implies that the KS HOMO energy is the opposite of the exact ionization energy

$$\varepsilon_{\mathsf{H}}^{\mathsf{N}} = -I_{\mathsf{N}}$$

▶ It is similar to Koopmans' theorem for HF, except that here it is exact (no neglect of correlation or orbital relaxation).

The LUMO energy, the electron affinity, the derivative discontinuity

Contrary to what one could have expected, the KS LUMO energy ε_L^N is not the opposite of the exact electron affinity $A_N = E_0^N - E_0^{N+1}$ but instead

$$\varepsilon_{\mathsf{L}}^{\mathsf{N}} = -A_{\mathsf{N}} - \Delta_{\mathsf{xc}}^{\mathsf{N}}$$

where $\Delta_{xc}^{N} > 0$ is a constant.

For the (N+1)-electron system (with the same external potential $v_{\rm ne}$), we have $\varepsilon_{\rm H}^{N+1}=-I_{N+1}=-A_N$, so it means that

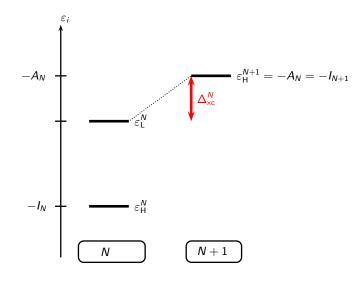
$$\Delta_{\mathsf{xc}}^{\mathit{N}} = arepsilon_{\mathsf{H}}^{\mathit{N}+1} - arepsilon_{\mathsf{L}}^{\mathit{N}}$$

- i.e., the constant Δ_{xc}^N corresponds to the "jump" of the LUMO energy of the N-electron system upon adding an electron so that the HOMO energy of the (N+1)-electron system correctly gives $-I_{N+1}$.
- In the extension of DFT to fractional electron numbers, it can be shown that the constant Δ_{xc}^N corresponds to the **uniform jump** that the **exchange-correlation potential** makes when going from $N-\delta$ electrons to $N+\delta$ electrons with $\delta \to 0^+$

$$\Delta_{\mathsf{xc}}^{N} = v_{\mathsf{xc}}^{N+\delta}(\mathbf{r}) - v_{\mathsf{xc}}^{N-\delta}(\mathbf{r}) = \left(\frac{\delta E_{\mathsf{xc}}[n]}{\delta n(\mathbf{r})}\right)_{N+\delta} - \left(\frac{\delta E_{\mathsf{xc}}[n]}{\delta n(\mathbf{r})}\right)_{N+\delta}$$

i.e. Δ_{xc}^N is the **derivative discontinuity** in the exchange-correlation energy functional $E_{xc}[n]$.

Kohn-Sham frontier orbital energies: Graphical summary



Fundamental gap

▶ The **fundamental gap** of the *N*-electron system is defined as

$$E_{\mathsf{gap}}^{N} = I_{N} - A_{N}$$

▶ In KS DFT, it thus be expressed as

$$E_{ ext{gap}}^N = \underbrace{arepsilon_{ ext{L}}^N - arepsilon_{ ext{H}}^N + \Delta_{ ext{xc}}^N}_{ ext{KS gap}} + \Delta_{ ext{xc}}^N$$

So the KS gap is not equal to the exact fundamental gap of the system, the difference coming from the derivative discontinuity Δ_{xc}^N .

▶ The derivative discontinuity Δ_{xc}^N can represent an important contribution to the fundamental gap. In the special case of open-shell systems, we have $\varepsilon_L^N = \varepsilon_H^N$, and thus if the fundamental gap of an open-shell system is not zero (Mott insulator), it is entirely given by Δ_{xc}^N .

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Local-density approximation

▶ In the **local-density approximation** (LDA), introduced by Kohn and Sham (1965), the exchange-correlation functional is approximated as

$$E_{xc}^{LDA}[n] = \int n(\mathbf{r}) \varepsilon_{xc}^{UEG}(n(\mathbf{r})) d\mathbf{r}$$

where $\varepsilon_{xc}^{UEG}(n)$ is the exchange-correlation energy per particle of the infinite **uniform** electron gas (UEG) with the density n.

▶ The exchange energy per particle of the UEG can be calculated analytically

$$\varepsilon_{\mathsf{x}}^{\mathsf{UEG}}(n) = C_{\mathsf{x}} \, n^{1/3}$$
 Dirac (1930) and Slater (1951)

For the correlation energy per particle $\varepsilon_c^{\text{UEG}}(n)$ of the UEG, there are some parametrized functions of n fitted to QMC data and imposing the high- and low-density expansions (using the Wigner-Seitz radius $r_s = (3/(4\pi n))^{1/3}$)

er-Seitz radius
$$r_s = (3/(4\pi n))^{1/3}$$
)
$$\varepsilon_c^{\text{UEG}} = A \ln r_s + B + C r_s \ln r_s + O(r_s) \qquad \text{high-density limit or weak-correlation limit}$$

$$\varepsilon_{\rm c}^{\rm UEG} = \frac{a}{r_{\rm s} \to \infty} + \frac{b}{r_{\rm s}^{3/2}} + \frac{c}{r_{\rm s}^2} + O\left(\frac{1}{r_{\rm s}^{5/2}}\right)$$
 low-density limit or strong-correlation limit

The two most used parametrizations are VWN and PW92. Generalization to spin densities $\varepsilon_c^{\text{UEG}}(n_\uparrow,n_\downarrow)$ is sometimes referred to as local-spin-density (LSD) approximation.

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The gradient-expansion approximation

- ► The next logical step beyond the LDA is the **gradient-expansion approximation** (GEA) which consists in a systematic expansion of $E_{xc}[n]$ in terms of the gradients of $n(\mathbf{r})$.
- ▶ To derive the GEA, one starts from the uniform electron gas, introduce a weak and slowly-varying external potential $v(\mathbf{r})$, and expand the exchange-correlation energy in terms of the gradients of the density. Alternatively, one can perform a semiclassical expansion of the exact $E_{\times c}[n]$.
- At second order, the GEA has the form

$$E_{\text{xc}}^{\text{GEA}}[n] = E_{\text{xc}}^{\text{LDA}}[n] + \int n(\mathbf{r})^{4/3} C_{\text{xc}}^{(2)}(n(\mathbf{r})) \left(\frac{\nabla n(\mathbf{r})}{n(\mathbf{r})^{4/3}}\right)^2 d\mathbf{r}$$

where $C_{xc}^{(2)}(n) = C_x^{(2)} + C_c^{(2)}(n)$ are known coefficients.

- lacktriangle We use the **reduced density gradient** $|\nabla n|/n^{4/3}$ which is a dimensionless quantity.
- ► The GEA should improve over the LDA provided that the reduced density gradient is small. Unfortunately, for real molecular systems, the reduced density gradient can be large in some regions of space, and the GEA turns out to be a worse approximation than the LDA.

Generalized-gradient approximations (1/4)

 The failure of the GEA lead to the development of generalized-gradient approximations (GGAs), started in the 1980s, of the generic form

$$E_{\mathsf{xc}}^{\mathsf{GGA}}[n] = \int e_{\mathsf{xc}}^{\mathsf{GGA}}(n(\mathbf{r}),
abla n(\mathbf{r})) d\mathbf{r}$$

- ▶ The GGAs provide a big improvement over LDA for molecular systems.
- ▶ The GGAs are often called **semilocal** approximations, which means that they involve a single integral on \mathbf{r} using "semilocal information" through $\nabla n(\mathbf{r})$.
- For simplicity, we consider here only the spin-independent form, but in practice GGA functionals are more generally formulated in terms of spin densities and their gradients

$$E_{xc}^{\mathsf{GGA}}[n] = \int e_{xc}^{\mathsf{GGA}}(n_{\uparrow}(\mathbf{r}), n_{\downarrow}(\mathbf{r}), \nabla n_{\uparrow}(\mathbf{r}), \nabla n_{\downarrow}(\mathbf{r})) d\mathbf{r}$$

 (Too) Many GGA functionals have been proposed. We will review some of the most widely used ones.

Generalized-gradient approximations (2/4)

▶ Becke 88 (B88 or B) exchange functional

$$E_{\mathbf{x}}^{\mathsf{B}}[n] = E_{\mathbf{x}}^{\mathsf{LDA}}[n] + \int n(\mathbf{r})^{4/3} f\left(\frac{|\nabla n(\mathbf{r})|}{n(\mathbf{r})^{4/3}}\right) d\mathbf{r}$$

- ▶ This form ensures the fulfilment of the scaling relation $E_x^B[n_\gamma] = \frac{\gamma}{\gamma} E_x^B[n]$.
- ► Function *f* chosen so as to satisfy the exact asymptotic behavior of the exchange energy per particle:

$$\varepsilon_{\mathsf{x}}(\mathbf{r}) \underset{r \to \infty}{\sim} -\frac{1}{2r}$$

- ▶ It contains an empirical parameter fitted to Hartree-Fock exchange energies of rare-gas atoms.
- ► Lee-Yang-Parr (LYP) correlation functional (1988)
 - ▶ One of the rare functionals not constructed starting from LDA.
 - ▶ It originates from the Colle-Salvetti (1975) correlation-energy approximation depending on the curvature of Hartree-Fock hole and containing four parameters fitted to Helium data.
 - LYP introduced a further approximation to retain dependence on only n, ∇n , $\nabla^2 n$.
 - ▶ The density Laplacian $\nabla^2 n$ can be exactly eliminated by an integration by parts.

Generalized-gradient approximations (3/4)

- ▶ Perdew-Wang 91 (PW91) exchange-correlation functional
 - ▶ It is based on a model of exchange and correlation holes from which we express the exchange and correlation energies per particle:

$$\varepsilon_{\mathsf{x}}(\mathbf{r}_1) = \frac{1}{2} \int \frac{n_{\mathsf{x}}(\mathbf{r}_1, \mathbf{r}_2)}{|\mathbf{r}_1 - \mathbf{r}_2|} \mathsf{d}\mathbf{r}_2 \quad \text{and} \quad \varepsilon_{\mathsf{c}}(\mathbf{r}_1) = \frac{1}{2} \int \frac{\bar{n}_{\mathsf{c}}(\mathbf{r}_1, \mathbf{r}_2)}{|\mathbf{r}_1 - \mathbf{r}_2|} \mathsf{d}\mathbf{r}_2$$

- ▶ It starts from the GEA model of these holes and removes the unrealistic long-range parts of these holes to restore important conditions satisfied by the LDA.
 - For the GEA exchange hole: the spurious positive parts are removed to enforce $n_x(\mathbf{r}_1, \mathbf{r}_2) \leq 0$ and a cutoff in $|\mathbf{r}_1 \mathbf{r}_2|$ is applied to enforce $\int n_x(\mathbf{r}_1, \mathbf{r}_2) d\mathbf{r}_2 = -1$.
 - ▶ For the GEA correlation hole: a cutoff is applied to enforce $\int \bar{n}_c(\mathbf{r}_1, \mathbf{r}_2) d\mathbf{r}_2 = 0$.
- ▶ The exchange and correlation energies per particle calculated from these numerical holes are then fitted to functions of n and $|\nabla n|$ chosen to satisfy a number of exact conditions.

Generalized-gradient approximations (4/4)

▶ Perdew-Burke-Ernzerhof (PBE) exchange-correlation functional (1996)

This is a simplification of the PW91 functional: $\varepsilon_{\rm x}$ and $\varepsilon_{\rm c}$ are simpler functions of n and $|\nabla n|$ enforcing fewer exact conditions and with no fitted parameters.

For exchange, the conditions imposed are:

- Second-order GEA expansion: $\varepsilon_x \underset{s\to 0}{\sim} \varepsilon_x^{\text{LDA}} + n^{1/3} C_x^{(2)} s^2$ where $s = |\nabla n|/n^{4/3}$ (with only approximate $C_x^{(2)} \approx -C_c^{(2)}$).
- ▶ A local version of Lieb-Oxford bound, $E_x \ge -C_{LO} \int n(\mathbf{r})^{4/3} d\mathbf{r}$. It was chosen to reach the bound in the $s \to \infty$ limit.

For correlation, the conditions imposed are:

- ▶ High-density limit: $\varepsilon_c \xrightarrow[r_s \to 0]{}$ const (cancellation of diverging term $A \ln r_s$ from LDA).
- ▶ Second-order GEA expansion: $\varepsilon_c \sim \varepsilon_c^{LDA} + n^{1/3} C_c^{(2)} s^2$ (with $C_c^{(2)}$ only in $r_s \to 0$ limit).
- ▶ Large reduced-density-gradient limit: $\varepsilon_c \xrightarrow[s \to \infty]{} 0$ (exchange dominates).

Meta-generalized-gradient approximations (1/2)

► The meta-generalized-gradient approximations (mGGAs) are of the generic form

$$E_{\mathsf{xc}}^{\mathsf{mGGA}}[n, au] = \int e_{\mathsf{xc}}^{\mathsf{mGGA}}(n(\mathbf{r}), \nabla n(\mathbf{r}), \nabla^2 n(\mathbf{r}), au(\mathbf{r})) d\mathbf{r}$$

where $\nabla^2 n(\mathbf{r})$ is the Laplacian of the density and $\tau(\mathbf{r})$ is the **non-interacting positive** kinetic energy density

$$au(\mathbf{r}) = rac{1}{2} \sum_{i=1}^{N} \left|
abla arphi_i(\mathbf{r})
ight|^2$$

which, as we will see, contains useful information.

- ▶ A mGGA can be either considered as an implicit functional of the density in the KS method, i.e. $E_{xc}[n] = E_{xc}^{mGGA}[n, \tau_{\Phi[n]}]$, or more commonly as an explicit functional of a single-determinant Φ in the GKS method, i.e. $E_{xc}^{S}[\Phi] = E_{xc}^{mGGA}[n_{\Phi}, \tau_{\Phi}]$.
- ▶ In the GKS method, a mGGA functional generates a non-multiplicative potential. But don't worry, this is allowed in GKS!
- Nowadays, $\nabla^2 n(\mathbf{r})$ is rarely used to construct mGGAs because it contains similar information than $\tau(\mathbf{r})$.
- ▶ The mGGAs are considered as part of the family of semilocal approximations.
- ► The mGGAs provide a modest improvement over GGAs.

Meta-generalized-gradient approximations (2/2)

- ▶ Motivations for introducing the variable $\tau(\mathbf{r})$:
 - ▶ Short-range expansion of the spherically average exchange hole (for closed-shell systems):

$$\tilde{n}_{x}(\mathbf{r}_{1}, r_{12}) = -\frac{n(\mathbf{r}_{1})}{2} - \frac{1}{3} \left(\frac{1}{4} \nabla^{2} n(\mathbf{r}_{1}) - 4\tau(\mathbf{r}_{1}) + \frac{|\nabla n(\mathbf{r}_{1})|^{2}}{8n(\mathbf{r}_{1})} \right) r_{12}^{2} + O(r_{12}^{4})$$

Thus $\tau(\mathbf{r})$ is needed to describe the curvature of the exchange hole.

 $\tau(\mathbf{r})$ can be used as an indicator of **one-orbital spatial regions** (regions containing one or two electrons in a single orbital). This is done by comparing $\tau(\mathbf{r})$ with the **von Weizsäcker kinetic energy density**

$$\tau^{\mathsf{W}}(\mathbf{r}) = \frac{|\nabla n(\mathbf{r})|^2}{8n(\mathbf{r})}$$

which is the exact $\tau(\mathbf{r})$ for one and two electrons in a single orbital.

- ▶ In practice, $\tau(\mathbf{r})$ is often used through the variables
 - $ightharpoonup z(\mathbf{r}) = \tau^{\mathsf{W}}(\mathbf{r})/\tau(\mathbf{r})$
 - $\alpha(\mathbf{r}) = (\tau(\mathbf{r}) \tau^{W}(\mathbf{r}))/\tau^{UEG}(\mathbf{r})$ where $\tau^{UEG}(\mathbf{r}) = c \ n(\mathbf{r})^{5/3}$
- Examples of mGGAs: TPSS (2003), M06-L (2006), and SCAN (2015).

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Hybrid approximations

In 1993, Becke proposed to mix Hartree-Fock (HF) exchange with GGA functionals in a three-parameter hybrid (3H) approximation $E_{xc}^{3H}[Φ] = a E_{xc}^{HF}[Φ] + b E_{xc}^{GGA}[n_Φ] + (1 - a - b) E_{xc}^{LDA}[n_Φ] + c E_{c}^{GGA}[n_Φ] + (1 - c) E_{c}^{LDA}[n_Φ]$

- Adding a fraction a of HF exchange decreases the self-interaction error, which tends to favor too much delocalized electron densities. However, a too large a tends to increase the static-correlation error (stretched chemical bonds, transition metal elements, ...).
- $E_{ ext{xc}}^{ ext{1H}}[\Phi] = a \; E_{ ext{x}}^{ ext{HF}}[\Phi] + (1-a) \; E_{ ext{x}}^{ ext{DFA}}[n_{\Phi}] + E_{ ext{c}}^{ ext{DFA}}[n_{\Phi}]$

In 1996, Becke proposed a simpler one-parameter hybrid (1H) approximation

- where E_x^{DFA} and E_c^{DFA} can be any semilocal density-functional approximations (DFAs).
- The optimal a is often around 0.25. Example: PBE0 = HF/PBE hybrid with a = 0.25.
 A strategy is to use flexible E_x^{DFA} and E_c^{DFA} in a hybrid approximation and optimize
 - many parameters on molecular properties. Example: B97 (13 parameters) and M06 and M06-2X (36 parameters).

Range-separated hybrid approximations

▶ Based on ideas of Savin (1996), Hirao and coworkers (2001) proposed a long-range correction (LC) scheme

$$E_{\mathsf{xc}}^{\mathsf{LC}}[\Phi] = E_{\mathsf{x}}^{\mathsf{Ir},\mathsf{HF}}[\Phi] + E_{\mathsf{x}}^{\mathsf{sr},\mathsf{DFA}}[n_{\Phi}] + E_{\mathsf{c}}^{\mathsf{DFA}}[n_{\Phi}]$$

where

- ► $E_x^{\text{Ir,HF}}[\Phi]$ is the HF exchange energy for the long-range electron-electron interaction $\frac{\text{erf}(\mu r_{12})}{r_{12}}$ replacing the Coulomb interaction $\frac{1}{r_{12}}$,
- $E_x^{sr,DFA}[n]$ is a semilocal DFA exchange energy for the complement **short-range electron-electron interaction** (semilocal DFAs are more accurate if limited to short-range interactions).
- ▶ the range-separation parameter μ (also sometimes denoted as ω) is often taken as $\mu \approx 0.3 0.5 \text{ bohr}^{-1}$.

Example: LC- ω PBE

► In 2004, Yanai, Tew, and Handy introduced a more flexible decomposition called the Coulomb-attenuating method (CAM)

$$E_{\text{xc}}^{\text{CAM}}[\Phi] = a \ E_{\text{x}}^{\text{sr,HF}}[\Phi] + b \ E_{\text{x}}^{\text{lr,HF}}[\Phi] + (1-a) \ E_{\text{x}}^{\text{sr,DFA}}[n_{\Phi}] + (1-b) \ E_{\text{x}}^{\text{lr,DFA}}[n_{\Phi}] + E_{\text{c}}^{\text{DFA}}[n_{\Phi}]$$

Examples: CAM-B3LYP, ω B97X

A special case: HSE (b = 0)

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Double-hybrid approximations

In 2006, Grimme introduced a two-parameter double-hybrid (2DH) approximation

$$E_{xc}^{2DH} = a_x E_x^{HF} [\Phi] + (1 - a_x) E_x^{DFA} [n_{\Phi}] + (1 - a_c) E_c^{DFA} [n_{\Phi}] + a_c E_c^{MP2}$$

where the MP2-like correlation energy E_c^{MP2} is added a posteriori with the previously calculated orbitals. Example: B2-PLYP ($a_x = 0.53$ and $a_c = 0.27$).

- The presence of nonlocal MP2 correlation allows one to use a larger fraction of nonlocal HF exchange.
- ▶ In 2011, Sharkas, Toulouse, and Savin showed that double hybrids can be understood as approximations of a **multideterminant extension of the KS method** based on the adiabatic-connection formalism in which the exact ground-state energy is written as

$$E_0 = \min_{\Psi \in \mathcal{W}} \left\{ \langle \Psi | \hat{T} + \hat{V}_{ne} + \lambda \hat{W}_{ee} | \Psi \rangle + \bar{E}_{Hxc}^{\lambda}[n_{\Psi}] \right\}$$
where $\bar{E}_{i}^{\lambda}[n] = (1 - \lambda)E_{i}[n] + (1 - \lambda)E[n] + \bar{E}_{i}^{\lambda}[n]$ and $\bar{E}_{i}^{\lambda}[n] = E[n] - \lambda^2 E_{i}[n_{\psi}]$

- where $\bar{E}_{\mathsf{Hxc}}^{\lambda}[n] = (1-\lambda)\mathsf{E}_{\mathsf{H}}[n] + (1-\lambda)\mathsf{E}_{\mathsf{x}}[n] + \bar{\mathsf{E}}_{\mathsf{c}}^{\lambda}[n]$ and $\bar{\mathsf{E}}_{\mathsf{c}}^{\lambda}[n] = \mathsf{E}_{\mathsf{c}}[n] \lambda^2 \mathsf{E}_{\mathsf{c}}[n_{1/\lambda}]$.
 - At second order of a non-linear Møller-Plesset-like perturbation theory, and using $E_c[n_{1/\lambda}] \approx E_c[n]$, we obtain a **one-parameter double-hybrid (1DH) approximation** $E_c^{1DH} = \lambda \ E_c^{HF}[\Phi] + (1-\lambda) \ E_c^{DFA}[n_{\Phi}] + (1-\lambda^2) E_c^{DFA}[n_{\Phi}] + \lambda^2 E_c^{MP2}$
 - ▶ The multideterminant extension of the KS method can also be used to rigorously combine wave-function methods such as MCSCF with DFT.

Range-separated double-hybrid approximations

► In 1996, Savin introduced the range-separated multideterminant extension of the KS scheme in which the exact ground-state energy is written as

$$\textit{E}_{0} = \min_{\Psi \in \mathcal{W}} \left\{ \langle \Psi | \hat{\textit{T}} + \hat{\textit{V}}_{\mathsf{ne}} + \hat{\textit{W}}_{\mathsf{ee}}^{\mathsf{Ir}} | \Psi \rangle + \bar{\textit{E}}_{\mathsf{Hxc}}^{\mathsf{sr}} [\textit{n}_{\Psi}] \right\}$$

where $\hat{W}^{\rm lr}_{\rm ee}$ is the long-range electron-electron operator for the pair potential ${\rm erf}(\mu r_{12})/r_{12}$ and $\bar{E}^{\rm sr}_{\rm Hxc}[n]$ is the complementary short-range density functional.

- ▶ The approach can be used to rigorously combine any wave-function method with DFT.
- ► In 2005, Ángyán, Gerber, Savin, and Toulouse introduced a range-separated double-hybrid (RSDH) approximation (also called RSH+MP2)

$$E_{xc}^{\mathsf{RSDH}} = E_{\mathsf{x}}^{\mathsf{lr},\mathsf{HF}}[\Phi] + E_{\mathsf{x}}^{\mathsf{sr},\mathsf{DFA}}[n_{\Phi}] + E_{\mathsf{c}}^{\mathsf{sr},\mathsf{DFA}}[n_{\Phi}] + E_{\mathsf{c}}^{\mathsf{lr},\mathsf{MP2}}$$

- ▶ Obtained as second order of a non-linear Møller-Plesset-like perturbation theory.
- ▶ Long-range MP2 is qualitatively correct for London dispersion interactions.
- ▶ Long-range MP2 has a fast convergence with the one-electron basis size.
- Extensions of this scheme to a more flexible CAM decomposition have also been proposed.

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Semiempirical dispersion corrections

 To explicitly account for London dispersion interactions, it has been proposed in the 2000s to add to the standard approximate functionals a semiempirical dispersion correction of the form

$$E_{ ext{disp}} = -s \sum_{lpha < eta} f(R_{lpha eta}) rac{C_6^{lpha eta}}{R_{lpha eta}^6}$$

where

- $ightharpoonup R_{\alpha\beta}$ is the distance between a pair of atoms,
- $ightharpoonup C_6^{\alpha\beta}$ is the dispersion coefficient between these atoms,
- $ightharpoonup f(R_{\alpha\beta})$ is a damping function which tends to 1 at large $R_{\alpha\beta}$ and tends to 0 at small $R_{\alpha\beta}$,
- ▶ s is a scaling parameter that can be adjusted for each approximate functional.
- ▶ The dispersion coefficients $C_6^{\alpha\beta}$ are empirically obtained from tabulated data.
- ▶ The most recent versions also includes $C_8^{\alpha\beta}$ two-body terms and $C_9^{\alpha\beta\gamma}$ three-body terms.
- ► This approach was named "DFT-D" by Grimme. Examples of DFT-D functionals:
- PBE-D, B97-D, B3LYP-D, ωB97X-D, B2PLYP-D.

 There are also various proposals to make the determination of dispersion coefficients less provided for Parks and Jahnson (2007). The banks and Sahafflay (2000). Saha and

empirical, e.g. Becke and Johnson (2007), Tkatchenko and Scheffler (2009), Sato and Nakai (2010).

Nonlocal van der Waals density functionals

► Another approach to describe dispersion interactions is to add to the standard approximate functionals a **nonlocal van der Waals density functional** of the form

$$E_{\mathrm{c}}^{\mathrm{nl}}[n] = \frac{1}{2} \iint n(\mathbf{r}_1) n(\mathbf{r}_2) \phi(\mathbf{r}_1, \mathbf{r}_2) \mathrm{d}\mathbf{r}_1 \mathrm{d}\mathbf{r}_2$$

where $\phi(\mathbf{r}_1, \mathbf{r}_2)$ is a correlation kernel.

- ➤ Two main families of such nonlocal correlation functionals exist: the "van der Waals density functionals" (vdW-DF) of Langreth, Lundqvist and coworkers and the Vydrov-Van Voorhis (VV) functionals.
- ▶ For example, the VV10 nonlocal correlation functional (2010) uses a theory-based kernel $\phi(\mathbf{r}_1, \mathbf{r}_2)$ with two adjustable parameters.
- Nonlocal van der Waals density functionals are less empirical but more computationally expensive than semiempirical dispersion corrections.
- \blacktriangleright Examples of functionals using VV10: ω B97X-V and ω B97M-V.

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 - Lieb-Oxford lower bound
 - Frontier orbital energies
- Usual approximations for the exchange-correlation energy
 - Local-density approximation
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- 4 Additional topics in density-functional theory
 - Time-dependent density-functional theory
 - Some less usual orbital-dependent exchange-correlation functionals
 - Fractional electron numbers and frontier orbital energies

Outline

Additional topics in density-functional theory

- Time-dependent density-functional theory
- Some less usual orbital-dependent exchange-correlation functionals
 Exact exchange
 Görling-Levy perturbation theory
 Adiabatic-connection fluctuation-dissipation approach
- Fractional electron numbers and frontier orbital energies
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Time-dependent density-functional theory (TDDFT)

Consider the time-dependent electronic Schrödinger equation with an external time-dependent potential $\hat{V}(t)$

$$irac{\partial |\Psi(t)
angle}{\partial t} = \left(\hat{T} + \hat{W}_{\mathsf{ee}} + \hat{V}(t)
ight) |\Psi(t)
angle$$

Similarly to the Hohenberg-Kohn theorem, **Runge and Gross** (1984) showed that, for a given initial wave function $\Psi(0)$, the time-dependent density $n(\mathbf{r},t)$ determines the time-dependent potential $v(\mathbf{r},t)$ up to an arbitrary additive time function:

$$n(\mathbf{r},t) \xrightarrow{\text{Runge-Gross}} v(\mathbf{r},t) + c(t)$$

► We can thus set up a time-dependent non-interacting KS system

$$i\frac{\partial \varphi_i(\mathbf{r},t)}{\partial t} = \left(-\frac{1}{2}\nabla^2 + \nu_{\rm s}(\mathbf{r},t)\right)\varphi_i(\mathbf{r},t)$$

where the time-dependent KS potential $v_{\rm s}({\bf r},t)=v({\bf r},t)+v_{\rm Hxc}({\bf r},t)$ reproduces the evolution of the exact density as $n({\bf r},t)=\sum_{i=1}^N|\varphi_i({\bf r},t)|^2$.

Remark: Runge and Gross also established a TDDFT variational theorem, but it was later shown to violate causality. Several different possible solutions to this problem have then been proposed.

Linear-response TDDFT

Let us consider a time-periodic potential of frequency ω . In Fourier space, a variation of the KS potential $v_s(\mathbf{r}_1,\omega)$ caused by a variation of the density $n(\mathbf{r}_2,\omega)$ can be written as

$$\frac{\delta v_{s}(\mathbf{r}_{1},\omega)}{\delta n(\mathbf{r}_{2},\omega)} = \frac{\delta v(\mathbf{r}_{1},\omega)}{\delta n(\mathbf{r}_{2},\omega)} + \frac{\delta v_{Hxc}(\mathbf{r}_{1},\omega)}{\delta n(\mathbf{r}_{2},\omega)}$$

This can be rewritten as

$$\chi_{\mathrm{s}}^{-1}(\mathbf{r}_1,\mathbf{r}_2,\omega) = \chi^{-1}(\mathbf{r}_1,\mathbf{r}_2,\omega) + f_{\mathrm{Hxc}}(\mathbf{r}_1,\mathbf{r}_2,\omega)$$

where

- ightarrow $\chi_{\rm s}({\bf r}_1,{\bf r}_2,\omega)=\delta {\it n}({\bf r}_1,\omega)/\delta {\it v}_{\rm s}({\bf r}_2,\omega)$ is the KS non-interacting linear-response function
- $ightharpoonup \chi(\mathbf{r}_1,\mathbf{r}_2,\omega)=\delta n(\mathbf{r}_1,\omega)/\delta v(\mathbf{r}_2,\omega)$ is the interacting linear-response function
- $f_{\rm Hxc}({\bf r}_1,{\bf r}_2,\omega)=\delta v_{\rm Hxc}({\bf r}_1,\omega)/\delta n({\bf r}_2,\omega)$ is the Hartree-exchange-correlation kernel
- ▶ The interacting linear-response function $\chi(\mathbf{r}_1,\mathbf{r}_2,\omega)$ is thus found from the **Dyson-like** response equation

$$\chi^{-1}(\mathbf{r}_1, \mathbf{r}_2, \omega) = \chi_s^{-1}(\mathbf{r}_1, \mathbf{r}_2, \omega) - f_{\mathsf{Hxc}}(\mathbf{r}_1, \mathbf{r}_2, \omega)$$

or, equivalently,

$$\chi(\mathbf{r}_1, \mathbf{r}_2, \omega) = \chi_{\mathrm{s}}(\mathbf{r}_1, \mathbf{r}_2, \omega) + \iint \mathrm{d}\mathbf{r}_3 \mathrm{d}\mathbf{r}_4 \; \chi_{\mathrm{s}}(\mathbf{r}_1, \mathbf{r}_3, \omega) \; f_{\mathrm{Hxc}}(\mathbf{r}_3, \mathbf{r}_4, \omega) \; \chi(\mathbf{r}_4, \mathbf{r}_2, \omega)$$

Excitation energies from linear-response TDDFT

▶ The KS linear-response function has poles at the KS (de-)excitation energies

$$\chi_{s}(\mathbf{r}_{1},\mathbf{r}_{2},\omega) = \sum_{\sigma \in \mathcal{S}} \sum_{i=1}^{\text{occ}} \sum_{j=1}^{\text{vir}} \left[\frac{\varphi_{i\sigma}^{*}(\mathbf{r}_{1})\varphi_{a\sigma}(\mathbf{r}_{1})\varphi_{a\sigma}^{*}(\mathbf{r}_{2})\varphi_{i\sigma}(\mathbf{r}_{2})}{\omega - (\varepsilon_{a} - \varepsilon_{i}) + i0^{+}} - \frac{\varphi_{a\sigma}^{*}(\mathbf{r}_{1})\varphi_{i\sigma}(\mathbf{r}_{1})\varphi_{i\sigma}^{*}(\mathbf{r}_{2})\varphi_{a\sigma}(\mathbf{r}_{2})}{\omega + (\varepsilon_{a} - \varepsilon_{i}) + i0^{+}} \right]$$

$$(\varepsilon_{a} - \varepsilon_{i}) + i0^{+} \qquad \omega + (\varepsilon_{a} - \varepsilon_{i}) + i0^{+} \qquad \omega$$

Similarly,
$$\chi(\mathbf{r}_1, \mathbf{r}_2, \omega)$$
 has poles at the exact excitation energies $\omega_n = E_n - E_0$.

$$\blacktriangleright \text{ Writing } \chi^{-1}(\omega) = \chi_{\text{s}}^{-1}(\omega) - f_{\text{Hxc}}(\omega) \text{ in the spin-orbital tensor-product basis } \{\psi_i^* \psi_a, \psi_a^* \psi_i\}$$

$$\boldsymbol{\chi}^{-1}(\omega) = -\left[\left(\begin{array}{cc} \mathbf{A}(\omega) & \mathbf{B}(\omega) \\ \mathbf{B}(-\omega)^* & \mathbf{A}(-\omega)^* \end{array}\right) - \omega \left(\begin{array}{cc} \mathbf{1} & \mathbf{0} \\ \mathbf{0} & -\mathbf{1} \end{array}\right)\right]$$

where the matrices $\mathbf{A}(\omega)$ and $\mathbf{B}(\omega)$ are

$$egin{align} [\mathbf{A}(\omega)]_{i\mathsf{a},jb} &= (arepsilon_{\mathsf{a}} - arepsilon_{i})\delta_{ij}\delta_{\mathsf{a}b} + \langle \mathsf{a}j|f_\mathsf{Hxc}(\omega)|ib
angle \ [\mathbf{B}(\omega)]_{i\mathsf{a},jb} &= \langle \mathsf{a}b|f_\mathsf{Hxc}(\omega)|ij
angle \end{aligned}$$

$$[\mathbf{D}(\omega)]_{la,jb} = \langle ab|_{\mathsf{Hxc}}(\omega)|y\rangle$$
The excitation energies ω_n can be calculated from the generalized eigenvalue equation

 $\begin{pmatrix} \mathbf{A}(\omega_n) & \mathbf{B}(\omega_n) \\ \mathbf{B}(-\omega_n)^* & \mathbf{A}(-\omega_n)^* \end{pmatrix} \begin{pmatrix} \mathbf{X}_n \\ \mathbf{Y}_n \end{pmatrix} = \omega_n \begin{pmatrix} \mathbf{1} & \mathbf{0} \\ \mathbf{0} & -\mathbf{1} \end{pmatrix} \begin{pmatrix} \mathbf{X}_n \\ \mathbf{Y}_n \end{pmatrix}$

The Hartree-exchange-correlation kernel

 In linear-response TDDFT, the key quantity to be approximated is the Hartree-exchange-correlation kernel

$$f_{\mathsf{Hxc}}(\mathbf{r}_1, \mathbf{r}_2, \omega) = \frac{\delta v_{\mathsf{Hxc}}(\mathbf{r}_1, \omega)}{\delta n(\mathbf{r}_2, \omega)}$$

It can be decomposed as

$$f_{\mathsf{Hxc}}(\mathbf{r}_1,\mathbf{r}_2,\omega) = f_{\mathsf{H}}(\mathbf{r}_1,\mathbf{r}_2) + f_{\mathsf{xc}}(\mathbf{r}_1,\mathbf{r}_2,\omega)$$

where the Hartree kernel is simply $f_H(\mathbf{r}_1, \mathbf{r}_2) = 1/|\mathbf{r}_1 - \mathbf{r}_2|$.

▶ In almost all TDDFT calculations, the frequency dependence of f_{xc} is neglected, which is called the **adiabatic approximation**

$$f_{\rm xc}({\bf r}_1,{\bf r}_2,\omega)\approx \frac{\delta v_{\rm xc}({\bf r}_1)}{\delta n({\bf r}_2)}=\frac{\delta E_{\rm xc}[n]}{\delta n({\bf r}_1)\delta n({\bf r}_2)}$$
 with the notorious consequence that only single-electron excitations are taken into

account (double excitations and higher are missing).

To describe nonlocal excitations, such as charge-transfer excitations, range-separated hybrid approximations are often used. The kernel has then the expression

$$f_{xc} = f_x^{lr,HF} + f_x^{sr,DFA} + f_c^{DFA}$$

where $f_x^{lr,HF}$ is the long-range HF exchange kernel.

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Orbital-dependent exchange-correlation functionals

- We discuss here some exchange-correlation energy functionals **depending explicitly on** the KS orbitals and KS orbital energies: $E_{xc}[\{\varphi_p\}, \{\varepsilon_p\}]$
- Since the KS orbitals and KS orbital energies are implicit functionals of the density, i.e. $\varphi_p[n](\mathbf{r})$ and $\varepsilon_p[n]$, these exchange-correlation expressions are **implicit functionals of** the density.
- ▶ In fact, the (range-separated) hybrid and double-hybrid approximations that we have seen already are sometimes considered to belong to this family, but they are more commonly considered within the GKS method, i.e. the orbitals are obtained with a nonlocal potential.
- Here, we are concerned with orbital-dependent exchange-correlation energy functionals within the KS method, i.e. with orbitals obtained with a **local potential**: $v_{xc}(\mathbf{r}) = \frac{\delta E_{xc}}{\delta n(\mathbf{r})}$
- ▶ Then, the calculation of the potential $v_{xc}(\mathbf{r})$ requires the **optimized-effective-potential** (OEP) method, which tends to be computationally involved.

Exact exchange

The exact exchange (EXX) energy functional is

$$\textit{E}_{\text{x}} = -\frac{1}{2} \sum_{\sigma \in \{\uparrow,\downarrow\}} \sum_{i=1}^{N_{\sigma}} \sum_{j=1}^{N_{\sigma}} \iint \frac{\varphi_{i\sigma}^{*}(\textbf{r}_{1})\varphi_{j\sigma}(\textbf{r}_{1})\varphi_{j\sigma}^{*}(\textbf{r}_{2})\varphi_{i\sigma}(\textbf{r}_{2})}{|\textbf{r}_{1} - \textbf{r}_{2}|} d\textbf{r}_{1} d\textbf{r}_{2}$$

It has exactly the same form as the HF exchange energy, but the orbitals used in this expression are different.

- The associated EXX potential $v_x(\mathbf{r})$ is calculated using the chain rule via the total KS potential $v_s(\mathbf{r})$
- $\frac{\delta E_{x}}{\delta v_{s}(\mathbf{r})} = \int \frac{\delta E_{x}}{\delta n(\mathbf{r}')} \frac{\delta n(\mathbf{r}')}{\delta v_{s}(\mathbf{r})} d\mathbf{r}'$
- Introducing the non-interacting KS static linear-response function $\chi_s(\mathbf{r}',\mathbf{r}) = \delta n(\mathbf{r}')/\delta v_s(\mathbf{r})$, we find the **OEP equation for the EXX potential**

$$\int v_{x}(\mathbf{r}')\chi_{s}(\mathbf{r}',\mathbf{r})d\mathbf{r}' = \frac{\delta E_{x}}{\delta v_{s}(\mathbf{r})}$$

Explicit expressions in terms of the orbitals can be derived for $\delta E_x/\delta v_s(\mathbf{r})$ and $\chi_s(\mathbf{r}',\mathbf{r})$.

The EXX occupied orbitals are very similar the HF ones, but the EXX virtual orbitals are much less diffuse than the HF ones $(v_x(\mathbf{r}) \sim -1/r)$ for all orbitals, contrary to HF).

Görling-Levy perturbation theory (1/2)

- In 1993, Görling and Levy developed a perturbation theory in terms of the coupling constant λ of the adiabatic connection.
- ▶ The Hamiltonian along the adiabatic connection (keeping the density constant) is

$$\hat{H}^{\lambda} = \hat{T} + \lambda \hat{W}_{\mathsf{ee}} + \hat{V}^{\lambda} = \hat{H}_{\mathsf{s}} + \lambda (\hat{W}_{\mathsf{ee}} - \hat{V}_{\mathsf{Hx}}) - \lambda^2 \hat{V}_{\mathsf{c}}^{(2)} - \cdots$$

where we have used $\hat{V}^{\lambda} = \hat{V}_{s} - \lambda \hat{V}_{Hx} - \hat{V}_{c}^{\lambda} = \hat{V}_{s} - \lambda \hat{V}_{Hx} - \lambda^{2} \hat{V}_{c}^{(2)} - \cdots$

lacktriangle At $\lambda=0$, we have the KS non-interacting reference Hamiltonian $\hat{H}_{s}=\hat{T}+\hat{V}_{s}$

$$\hat{H}_{\mathsf{s}}|\Phi_n\rangle=\mathcal{E}_n|\Phi_n\rangle$$

where $\Phi_0 \equiv \Phi$ is the ground-state KS determinant.

lacktriangle The ground-state wave function Ψ^λ of \hat{H}^λ is expanded in powers of λ

$$|\Psi^{\lambda}
angle = |\Phi
angle + \lambda |\Psi^{(1)}
angle + \cdots$$
 with $|\Psi^{(1)}
angle = -\sum_{n
eq 0} rac{\langle \Phi_n | \hat{W}_{\text{ee}} - \hat{V}_{\text{Hx}} |\Phi
angle}{\mathcal{E}_n - \mathcal{E}_0} |\Phi_n
angle$

assuming a nondegenerate KS ground state.

Görling-Levy perturbation theory (2/2)

 \blacktriangleright The correlation energy is then expanded in powers of λ

$$E_{\mathsf{c}}^{\lambda} = \langle \Psi^{\lambda} | \hat{T} + \lambda \hat{W}_{\mathsf{ee}} | \Psi^{\lambda} \rangle - \langle \Phi | \hat{T} + \lambda \hat{W}_{\mathsf{ee}} | \Phi \rangle = E_{\mathsf{c}}^{(0)} + \lambda E_{\mathsf{c}}^{(1)} + \lambda^2 E_{\mathsf{c}}^{(2)} + \cdots$$

where the zeroth- and first-order terms vanish: $E_{\rm c}^{(0)}=0$ and $E_{\rm c}^{(1)}=(\partial E_{\rm c}^{\lambda}/\partial \lambda)_{\lambda=0}=0$

The second-order term is the **second-order Görling-Levy (GL2) correlation energy**
$$E_c^{\text{GL2}} \equiv E_c^{(2)} = \langle \Phi | \hat{W}_{\text{ee}} | \Psi^{(1)} \rangle = \langle \Phi | \hat{W}_{\text{ee}} - \hat{V}_{\text{Hx}} | \Psi^{(1)} \rangle$$

where we have used $\langle \Phi | \hat{V}_{\text{Hx}} | \Psi^{(1)} \rangle = 0$ since it is the derivative with respect to λ at $\lambda = 0$ of $\langle \Psi^{\lambda} | \hat{V}_{\text{Hx}} | \Psi^{\lambda} \rangle = \int v_{\text{Hx}}(\mathbf{r}) n(\mathbf{r}) d\mathbf{r}$ which does not depend on λ .

► The GL2 correlation energy is thus

$$E_{\rm c}^{\sf GL2} = -\sum_{\sf r=0} rac{|\langle \Phi | \hat{W}_{\sf ee} - \hat{V}_{\sf Hx} | \Phi_n
angle|^2}{\mathcal{E}_n - \mathcal{E}_0} = E_{\sf c}^{\sf MP2} + E_{\sf c}^{\sf S}$$

with a double-excitation MP2-like term E_c^{MP2} and single-excitation term E_c^{S}

$$E_{c}^{\text{MP2}} = -\frac{1}{4} \sum_{i,j}^{\text{occ}} \sum_{a,b}^{\text{vir}} \frac{|\langle ij||ab\rangle|^{2}}{\varepsilon_{a} + \varepsilon_{b} - \varepsilon_{i} - \varepsilon_{j}} \text{ and } E_{c}^{\text{S}} = -\sum_{i}^{\text{occ}} \sum_{a}^{\text{vir}} \frac{|\langle i|\hat{V}_{x}^{\text{HF}} - \hat{V}_{x}|a\rangle|^{2}}{\varepsilon_{a} - \varepsilon_{i}}$$

▶ In practice, results are often disappointing! It is preferable to go beyond second order with the random-phase approximation.

Adiabatic-connection fluctuation-dissipation approach (1/2)

▶ The adiabatic-connection formula for the correlation energy is

$$\textit{E}_{c} = \frac{1}{2} \int_{0}^{1} \mathsf{d}\lambda \iint \frac{n_{2,c}^{\lambda}(\textbf{r}_{1},\textbf{r}_{2})}{|\textbf{r}_{1}-\textbf{r}_{2}|} \mathsf{d}\textbf{r}_{1} \mathsf{d}\textbf{r}_{2}$$

► The correlation part of the pair density can be written

$$n_{2,c}^{\lambda}(\mathbf{r}_1,\mathbf{r}_2) = \langle \Psi^{\lambda} | \hat{n}_2(\mathbf{r}_1,\mathbf{r}_2) | \Psi^{\lambda} \rangle - \langle \Phi | \hat{n}_2(\mathbf{r}_1,\mathbf{r}_2) | \Phi \rangle$$

where $\hat{n}_2(\mathbf{r}_1, \mathbf{r}_2)$ is the pair-density operator.

We use the expression of the pair-density operator in terms of the density operator $\hat{n}(\mathbf{r})$

$$\hat{n}_2(\mathbf{r}_1,\mathbf{r}_2) = \hat{n}(\mathbf{r}_1)\hat{n}(\mathbf{r}_2) - \delta(\mathbf{r}_1 - \mathbf{r}_2)\hat{n}(\mathbf{r}_1)$$

and the fact that the density is constant along the adiabatic connection

$$\langle \Psi^{\lambda} | \hat{n}(\mathbf{r}_1) | \Psi^{\lambda}
angle = \langle \Phi | \hat{n}(\mathbf{r}_1) | \Phi
angle$$

▶ We thus see that the correlation pair density can be written as

$$\textit{n}^{\lambda}_{2,c}(\textbf{r}_{1},\textbf{r}_{2}) = \langle \Psi^{\lambda} | \hat{\textit{n}}(\textbf{r}_{1}) \hat{\textit{n}}(\textbf{r}_{2}) | \Psi^{\lambda} \rangle - \langle \Phi | \hat{\textit{n}}(\textbf{r}_{1}) \hat{\textit{n}}(\textbf{r}_{2}) | \Phi \rangle$$

Adiabatic-connection fluctuation-dissipation approach (2/2)

▶ Let us consider the linear-response function along the adiabatic connection

$$\chi_{\lambda}(\mathbf{r}_{1},\mathbf{r}_{2},\omega) = \sum_{n\neq0} \frac{\langle \Psi^{\lambda} | \hat{n}(\mathbf{r}_{1}) | \Psi^{\lambda}_{n} \rangle \langle \Psi^{\lambda}_{n} | \hat{n}(\mathbf{r}_{2}) | \Psi^{\lambda} \rangle}{\omega - \omega^{\lambda}_{n} + i0^{+}} - \frac{\langle \Psi^{\lambda} | \hat{n}(\mathbf{r}_{2}) | \Psi^{\lambda}_{n} \rangle \langle \Psi^{\lambda}_{n} | \hat{n}(\mathbf{r}_{1}) | \Psi^{\lambda} \rangle}{\omega + \omega^{\lambda}_{n} + i0^{+}}$$

where the sum is over all eigenstates Ψ_n^{λ} of the Hamiltonian \hat{H}^{λ} , i.e. $\hat{H}^{\lambda}|\Psi_n^{\lambda}\rangle = E_n^{\lambda}|\Psi_n^{\lambda}\rangle$, except the ground state $\Psi^{\lambda} \equiv \Psi_0^{\lambda}$, and $\omega_n^{\lambda} = E_n^{\lambda} - E_0^{\lambda}$ are the excitation energies.

> By contour integrating $\chi_{\lambda}(\mathbf{r}_1,\mathbf{r}_2,\omega)$ around the right half ω-complex plane, we arrive at the **fluctuation-dissipation theorem**

$$n_{2,c}^{\lambda}(\mathbf{r}_1,\mathbf{r}_2) = -\int_{-\infty}^{+\infty} \frac{\mathrm{d}\omega}{2\pi} \left[\chi_{\lambda}(\mathbf{r}_1,\mathbf{r}_2,i\omega) - \chi_{\mathrm{s}}(\mathbf{r}_1,\mathbf{r}_2,i\omega) \right]$$

which relates ground-state correlations in the time-independent system $n_{2,c}^{\lambda}(\mathbf{r}_1,\mathbf{r}_2)$ to the linear response of the system $\chi_{\lambda}(\mathbf{r}_1,\mathbf{r}_2,\omega)$ due to a time-dependent external perturbation.

► We thus have the adiabatic-connection fluctuation-dissipation (ACFD) formula for the correlation energy

$$\textit{E}_{c} = -\frac{1}{2} \int_{0}^{1} d\lambda \int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} \iint \frac{\chi_{\lambda}(\textbf{r}_{1},\textbf{r}_{2},i\omega) - \chi_{s}(\textbf{r}_{1},\textbf{r}_{2},i\omega)}{|\textbf{r}_{1}-\textbf{r}_{2}|} d\textbf{r}_{1} d\textbf{r}_{2}$$

Random-phase approximation

▶ The ACFD formula involves $\chi_{\lambda}(\mathbf{r}_1,\mathbf{r}_2,i\omega)$ which can be obtained from linear-response TDDFT

$$\chi_{\lambda}(\mathbf{r}_{1},\mathbf{r}_{2},\omega) = \chi_{s}(\mathbf{r}_{1},\mathbf{r}_{2},\omega) + \iint d\mathbf{r}_{3}d\mathbf{r}_{4} \ \chi_{s}(\mathbf{r}_{1},\mathbf{r}_{3},\omega) \ f_{\mathsf{Hxc}}^{\lambda}(\mathbf{r}_{3},\mathbf{r}_{4},\omega) \ \chi_{\lambda}(\mathbf{r}_{4},\mathbf{r}_{2},\omega)$$

The simplest approximation is the (direct) random-phase approximation (RPA)

▶ By iterating the TDDFT response equation, we find the RPA linear-response function

 $f_{\mathsf{Hyc}}^{\lambda}(\mathbf{r}_1,\mathbf{r}_2,\omega)\approx f_{\mathsf{H}}^{\lambda}(\mathbf{r}_1,\mathbf{r}_2)=\lambda w_{\mathsf{ee}}(\mathbf{r}_1,\mathbf{r}_2)$

where
$$w_{\rm ee}({\bf r}_1,{\bf r}_2)=1/|{\bf r}_1-{\bf r}_2|$$
 is the Coulomb interaction.

 $\chi_{\lambda}^{\text{RPA}}(\omega) = \chi_{\text{s}}(\omega) + \lambda \chi_{\text{s}}(\omega) \mathbf{w}_{\text{ee}} \chi_{\text{s}}(\omega) + \lambda^2 \chi_{\text{s}}(\omega) \mathbf{w}_{\text{ee}} \chi_{\text{s}}(\omega) \mathbf{w}_{\text{ee}} \chi_{\text{s}}(\omega) + \cdots$

Finally, the (direct) RPA correlation energy is

 $E_{\rm c}^{\rm RPA} = -\frac{1}{2} \int_{\rm a}^{1} {\rm d}\lambda \int_{\rm a}^{+\infty} \frac{{\rm d}\omega}{2\pi} {\rm Tr} \Big[{\bf w}_{\rm ee} \Big(\lambda {\bf \chi}_{\rm s}(i\omega) {\bf w}_{\rm ee} {\bf \chi}_{\rm s}(i\omega) + \lambda^2 {\bf \chi}_{\rm s}(i\omega) {\bf w}_{\rm ee} {\bf \chi}_{\rm s}(i\omega) {\bf w}_{\rm ee} {\bf \chi}_{\rm s}(i\omega) + \cdots \Big) \Big]$

▶ The (direct) RPA correlation energy corresponds to the sum of all direct ring diagrams. It accounts for long-range van der Waals dispersion interactions. However, it shows large self-interaction errors, which can be overcome by adding exchange diagrams.

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Quantum mechanics with fractional electron numbers (1/2)

The ground-state energy of a system with a **fractional number of electrons** $\mathcal{N} = \mathbb{N} - 1 + f$ (where \mathbb{N} is an integer and $0 \le f \le 1$) can be defined as

$$E_0^{N-1+f} = \min_{\hat{\Gamma} \in \mathcal{D}_{\mathsf{DM}}^{N-1+f}} \mathsf{Tr} \left[\left(\hat{T} + \hat{W}_{\mathsf{ee}} + \hat{V}_{\mathsf{ne}} \right) \hat{\Gamma} \right]$$

where the minimization is over **ensemble density matrices** $\hat{\Gamma}$ in the set

$$\mathcal{D}_{\mathsf{DM}}^{N-1+f} = \left\{ \hat{\Gamma} = (1-f)|\Psi^{N-1}\rangle\langle\Psi^{N-1}| + f|\Psi^{N}\rangle\langle\Psi^{N}|, \ \Psi^{N-1} \in \mathcal{W}^{N-1}, \Psi^{N} \in \mathcal{W}^{N} \right\}$$

where f is fixed, and Ψ^{N-1} and Ψ^N are arbitrary wave functions in the (N-1)- and N-electron admissible wave-function sets \mathcal{W}^{N-1} and \mathcal{W}^N , respectively.

$$\hat{\Gamma}_0 = (1-f) |\Psi_0^{N-1}
angle \langle \Psi_0^{N-1}| + f |\Psi_0^N
angle \langle \Psi_0^N|$$

where Ψ_0^{N-1} and Ψ_0^N are the (N-1)- and N-electron ground-state wave functions.

 $E_0^{N+f} = (1-f) E_0^N + f E_0^{N+1}$

▶ The ground-state energy is **linear** in f between the integer numbers N-1 and N

$$E_0^{N-1+f} = (1-f) E_0^{N-1} + f E_0^N$$

where E_0^{N-1} and E_0^N are the (N-1)- and N-electron ground-state energies.

Similarly, between the integer electron numbers N and N+1, we have

Quantum mechanics with fractional electron numbers (2/2)

► The ground-state energy is a **continuous piecewise**

respect to f, we find for $N-1 < \mathcal{N} < N$

 $\left(\frac{\partial E_0^{\mathcal{N}}}{\partial \mathcal{N}}\right)_{N-1<\mathcal{N}< N} = E_0^N - E_0^{N-1} = -I_N$ where I_N is the **ionization energy** of the N-electron system.

linear function of the fractional electron number
$$\mathcal{N}$$
.

The derivative of $E_0^{\mathcal{N}}$ with respect to \mathcal{N} defines the electronic chemical potential

$$\mu = \frac{\partial E_0^{\mathcal{N}}}{\partial \mathcal{N}}$$

Taking the derivative with respect to \mathcal{N} corresponds to taking the derivative with

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Similarly for
$$N<\mathcal{N}< N+1$$

$$\left(\frac{\partial E_0^\mathcal{N}}{\partial \mathcal{N}}\right)_{N<\mathcal{N}< N+1} = E_0^{N+1} - E_0^N = -A_N$$

where A_N is the **electron affinity** of the N-electron system. The electronic chemical potential μ has thus a **discontinuity** at the integer electron number N. So, the plot of E_0^N with respect to N is made of a series of straight lines between integer electron numbers, with derivative discontinuities at each integer.

DFT with fractional electron numbers (1/2)

▶ The universal density functional F[n] is extended to densities integrating to a fractional electron number, $\int n(\mathbf{r})d\mathbf{r} = \mathcal{N} = N - 1 + f$, as

$$F[n] = \min_{\substack{\hat{\Gamma} \in \mathcal{D}_{\mathrm{DM}}^{N-1+f} \\ \hat{\Gamma} \to n}} \mathrm{Tr}\left[\left(\hat{T} + \hat{W}_{\mathrm{ee}} \right) \hat{\Gamma} \right]$$

▶ To set up a KS method, we introduce the decomposition $F[n] = T_s[n] + E_{Hxc}[n]$ where

$$T_s[n] = \min_{\substack{\hat{\Gamma}_s \in \mathcal{D}_{\mathrm{DM},s}^{N-1+f} \\ \hat{\Gamma}_s \to n}} \mathrm{Tr}[\hat{\mathcal{T}}\hat{\Gamma}_s]$$
 is the KS non-interacting kinetic-energy functional and the minimiz

is the KS non-interacting kinetic-energy functional and the minimization is over ensemble non-interacting density matrices $\hat{\Gamma}_s$ in the set

$$\mathcal{D}_{\mathsf{DM,s}}^{N-1+f} = \left\{ \hat{\mathsf{\Gamma}}_{\mathsf{s}} = (1-f) |\Phi^{N-1,f}\rangle \langle \Phi^{N-1,f}| + f |\Phi^{N,f}\rangle \langle \Phi^{N,f}| \right\}$$

where $\Phi^{N-1,f}$ and $\Phi^{N,f}$ are (N-1)- and N-electron single-determinant wave functions, constructed from a **common set of orbitals** $\{\varphi_i\}$ depending on the fixed f.

The exact ground-state energy can then be expressed as -N-1+f -N-1+f -N-1+f -N-1+f

$$E_0^{N-1+f} = \min_{\hat{\Gamma}_s \in \mathcal{D}_{\mathrm{DM},s}^{N-1+f}} \Bigl\{ \mathsf{Tr} \left[\left(\hat{\mathcal{T}} + \hat{\mathcal{V}}_{\mathsf{ne}} \right) \hat{\Gamma}_s \right] + E_{\mathsf{Hxc}}[n_{\hat{\Gamma}_s}] \Bigr\}$$

DFT with fractional electron numbers (2/2)

The total energy can be written in terms of **orbital occupation numbers** n_i

$$E^{N-1+f} = \sum_{i=1}^{N} n_i \int \varphi_i^*(\mathbf{r}) \left(-rac{1}{2}
abla^2 +
u_{\mathsf{ne}}(\mathbf{r})
ight) arphi_i(\mathbf{r}) \mathrm{d}\mathbf{r} + E_{\mathsf{Hxc}}[n]$$

with the density $n(\mathbf{r}) = \sum_{i=1}^{N} n_i |\varphi_i(\mathbf{r})|^2$ and the occupation numbers $n_i = 1$ for $i \leq N-1$ and $n_N = \mathbf{f}$ for the HOMO (ignoring degeneracy for simplicity).

▶ The orbitals satisfy standard-looking KS equations

$$\left(-\frac{1}{2}\nabla^2 + \nu_{\rm s}({\bf r})\right)\varphi_i({\bf r}) = \varepsilon_i\varphi_i({\bf r}) \quad \text{with} \quad \nu_{\rm s}({\bf r}) = \nu_{\rm ne}({\bf r}) + \frac{\delta E_{\rm Hxc}[n]}{\delta n({\bf r})}$$
 with the important difference that we can now fix the arbitrary constant in $\nu_{\rm s}({\bf r})$.

This is because we can now allow variations of $n(\mathbf{r})$ changing \mathcal{N} , i.e. $\int \delta n(\mathbf{r}) d\mathbf{r} \neq 0$

$$\delta E_{\mathsf{Hxc}}[n] = \int \left(\frac{\delta E_{\mathsf{Hxc}}[n]}{\delta n(\mathbf{r})} + \mathsf{const} \right) \delta n(\mathbf{r}) d\mathbf{r}$$

making the constant no longer arbitrary. This unambiguously fixes the values of the KS orbital energies ε_i .

Janak's theorem (1978): After optimizing the orbitals with fixed occupation numbers, we have

$$\frac{\partial E^N}{\partial \mathbf{r}_i} = \varepsilon_i$$
 for occupied orbitals

The HOMO energy and the ionization energy

- For clarity in the discussion, we will now explicitly indicate the dependence on the electron number \mathcal{N} .
- ▶ Janak's theorem applied to the HOMO for $\mathcal{N} = N \delta$ where $\delta \to 0^+$ gives

$$\left(\frac{\partial E_0^{\mathcal{N}}}{\partial \mathcal{N}}\right)_{N-\delta} = \varepsilon_{\mathsf{H}}^{N-\delta} \equiv \varepsilon_{\mathsf{H}}^{N}$$

where $\varepsilon_{\rm H}^{\it N}$ is the HOMO energy of the *N*-electron system (defined as the left side of discontinuity).

► This implies that the KS HOMO energy is the opposite of the exact ionization energy

$$\varepsilon_{\mathsf{H}}^{\mathsf{N}} = -I_{\mathsf{N}}$$

► Combining this result with the known asymptotic behavior of the exact ground-state density (for finite systems)

$$n^N(\mathbf{r}) \sim e^{-2\sqrt{2I_N} r}$$

it can be shown that it implies that the KS potential $v_s^N(\mathbf{r}) \equiv v_s^{N-\delta}(\mathbf{r})$ (defined as the limit from the left side) vanishes asymptotically

$$\lim_{r\to\infty} v_{\rm s}^N(\mathbf{r}) = 0$$

Janak's theorem applied to the HOMO but now for $\mathcal{N} = N + \delta$ where $\delta \to 0^+$ gives $\left(\frac{\partial E_0^N}{\partial \mathcal{N}}\right)_{N+\delta} = \varepsilon_{\mathsf{H}}^{N+\delta} = -A_N$

The LUMO energy, the electron affinity, the derivative discontinuity (1/2)

where
$$\varepsilon_{\rm H}^{N+\delta}$$
 is the HOMO energy from the right side of the discontinuity.

Remark: $\partial E_0^{\mathcal{N}}/\partial \mathcal{N}$ is constant for all $N < \mathcal{N} < N+1$, so: $\varepsilon_H^{N+\delta} = \varepsilon_H^{N+1-\delta} \equiv \varepsilon_H^{N+1}$

One may think that
$$\varepsilon_{\rm H}^{N+\delta}$$
 is equal to the LUMO energy of the N-electron system $\varepsilon_{\rm L}^N$

$$\varepsilon_{\rm H}^{N+\delta} \stackrel{?}{\not{=}} \varepsilon_{\rm L}^N \equiv \varepsilon_{\rm L}^{N-\delta} \qquad \qquad \text{but this is WRON}$$

Let us compare $\varepsilon_{H}^{N+\delta}$ and $\varepsilon_{L}^{N-\delta}$:

s compare
$$\varepsilon_{\mathrm{H}}^{N+\delta}$$
 and $\varepsilon_{\mathrm{L}}^{N-\delta}$:
$$\varepsilon_{\mathrm{H}}^{N+\delta} = \int \varphi_{\mathrm{H}}^{N+\delta}(\mathbf{r})^* \left(-\frac{1}{2}\nabla^2 + v_{\mathrm{s}}^{N+\delta}(\mathbf{r})\right) \varphi_{\mathrm{H}}^{N+\delta}(\mathbf{r}) \mathrm{d}\mathbf{r}$$

$$\varepsilon_{\mathrm{L}}^{N-\delta} = \int \varphi_{\mathrm{L}}^{N-\delta}(\mathbf{r})^* \left(-\frac{1}{2}\nabla^2 + v_{\mathrm{s}}^{N-\delta}(\mathbf{r})\right) \varphi_{\mathrm{L}}^{N-\delta}(\mathbf{r}) \mathrm{d}\mathbf{r}$$

The density is continuous at the integer N, i.e. $n^{N+\delta}(\mathbf{r}) = n^{N-\delta}(\mathbf{r})$, but this only imposes that $v_s^{N+\delta}(\mathbf{r})$ and $v_s^{N-\delta}(\mathbf{r})$ be equal up to an additive constant (according to the Hohenberg-Kohn theorem).

but this is WRONG!

The LUMO energy, the electron affinity, the derivative discontinuity (2/2) • Indeed, it turns out that $v_s^{N+\delta}(\mathbf{r})$ and $v_s^{N-\delta}(\mathbf{r})$ do differ by a uniform constant Δ_{xc}^N

$$v_{\mathsf{s}}^{N+\delta}(\mathbf{r}) - v_{\mathsf{s}}^{N-\delta}(\mathbf{r}) = \Delta_{\mathsf{xc}}^{N}$$

The orbitals are continuous at the integer N, so $\varphi_H^{N+\delta}(\mathbf{r}) = \varphi_L^{N-\delta}(\mathbf{r})$, and we find

$$arepsilon_{\mathsf{H}}^{\mathit{N}+\delta} = arepsilon_{\mathsf{L}}^{\mathit{N}-\delta} + \Delta_{\mathsf{xc}}^{\mathit{N}}$$

In conclusion, the KS LUMO energy is not the opposite of the exact electron affinity $\varepsilon_1^N = -A_N - \Delta_{\rm ex}^N$

due to the discontinuity Δ_{xc}^{N} in the KS potential.

Such a discontinuity can only come from the exchange-correlation part of the potential $v_{xc}^{\mathcal{N}}(\mathbf{r})$ since $v_{ne}(\mathbf{r})$ is independent from \mathcal{N} and the Hartree potential

$$v_{\rm H}^{\mathcal{N}}(\mathbf{r}) = \int n^{\mathcal{N}}(\mathbf{r}')/|\mathbf{r} - \mathbf{r}'| \mathrm{d}\mathbf{r}' \text{ is a continuous function of } \mathcal{N}. \text{ So, we have}$$

$$\Delta_{\rm xc}^{\mathcal{N}} = v_{\rm xc}^{\mathcal{N}+\delta}(\mathbf{r}) - v_{\rm xc}^{\mathcal{N}-\delta}(\mathbf{r}) = \left(\frac{\delta E_{\rm xc}[n]}{\delta n(\mathbf{r})}\right)_{\rm yc.s.} - \left(\frac{\delta E_{\rm xc}[n]}{\delta n(\mathbf{r})}\right)_{\rm yc.s.}$$

i.e. Δ_{xc}^{N} is the **derivative discontinuity** in the exchange-correlation energy functional $E_{xc}[n]$.

Kohn-Sham frontier orbital energies: Graphical summary

