Problème quantique à N-corps en physique nucléaire

A journey through a general introduction and an overview of the « ab initio » approach



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• Introduction to low-energy nuclear physics

 \circ Phenomenology

Rationale from the theoretical viewpoint

• Strong inter-nucleon forces

• Phenomenology and modern modelling

• The ab initio nuclear many-body problem

• Specificities

Recent developments

• Status of ab initio calculations of medium-mass atomic nuclei

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Elementary facts and « big » questions about nuclei

252 stable isotopes, ~3100 synthesized in the lab
How many bound (w.r.t strong force) nuclei exist; 9000?



Updated in 2019 to Z=9 (22 neutrons) and Z=10 (24 neutrons)

• Neutron **drip-line** known up to $\frac{Z-8}{16}$ neutrons) • Where is the neutron drip-line beyond Z=10? Oganesson 118Og added to Mendeleïev table in 2016 • Heaviest synthesized element Z=118 • Heaviest possible element? • Enhanced stability near Z=120?126?

2p decay beyond the proton drip line in ⁴⁵Fe in 2002
O Modes of instability (α, p, β, γ decays, fission)
Are there more exotic/rare decay modes?
Ex: v-less 2β decay = test of standard model?

Gravitational wave + kilonova from neutron stars merger in 2017
Elements up to Fe produced in stellar fusion
How have heavier elements been produced?
Exotic r-process nucleosynthesis ; but where?

Shown to disappear away from stability in 1975/1993

Over-stable "magic" nuclei (2, 8, 20, 28, 50, 82, ...)
How other magic numbers evolve with N-Z?

The atomic nucleus as a 4-components quantum mesoscopic system *An extremely rich and diverse phenomenology*



The atomic nucleus as a laboratory test *A multi-scale and multi-physics connector*



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Energy scales and degrees of freedom

Emergent phenomena

amenable

to effective descriptions



High-energy nuclear, i.e. hadronic, physics

- Realm of Quantum Chromo Dynamics
- Quarks and gluons
- Chiral symmetry of QCD
- Confinement and asymptotic freedom

The description

- Depends on the energy scale
- > Must rely on appropriate choice of DOFs
- Must encode the key symmetries

Low-energy nuclear physics

- > Of order of ~150MeV (M_{nuc} , m_{π} ...)
- Nucleons and pions
- > Chiral symmetry (breaking) of QCD
- Even more effective DOFs for MeV scale

Nuclear physics moving from a plurality of nuclear models...



...to an arborescence of nuclear effective (field) theories



Ab initio (i.e. In medias res) quantum many-body problem

Ab initio ("from scratch") scheme = A-body Schrödinger Equation (SE)

$$H|\Psi_k^{\rm A}\rangle = E_k^{\rm A}|\Psi_k^{\rm A}\rangle$$

A-body Hamiltonian

$$H = T + V^{2N} + V^{3N} + V^{4N} + \ldots + V^{AN}$$



A structure-less nucleons as d.o.f
All nucleons active in complete Hilbert space
Elementary interactions between them
Solve A-body Schroedinger equation (SE)

• Thorough estimate of error



Do we know the form of V^{2N}, V^{3N} etc **Do we know how to derive them from QCD?** Why would there be forces beyond pairwise? **Consistent construction of other operators?**



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Phenomenology of inter-nucleon interactions



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a. Strong central + spin-orbit + tensor operators
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- b. Dominant two-nucleon and sub-leading (but mandatory) three-nucleon operators
- 3. Infra-red source of non-perturbativeness

a. Large scattering lengthes (near unitarity) in s-wave channels = nn virtual state/np bound state

S=0

T=1

S=1

T=0

- 4. Ultra-violet source of non-perturbativeness
 - c. Short-range repulsion

Modern constructive approach = effective field theory

 Use separation of scales to define d.o.f & expansion parameter [Weinberg, Gasser, Leutwyler, van Kolck, ..] Typical momentum at play → Q/Λ → High energy scale (physics beyond not included explicitly)
 Parametrize physics beyond Λ + write #∞ terms allowed by (broken) symmetries of underlying QCD
 Order by size all possible terms → systematic expansion ("power counting") → theoretical error
 Truncate at a given order and adjust low-energy constants (LECs) via underlying theory or data
 Regularize UV divergences and (hopefully) achieve order-by-order renormalization of observables



Chiral effective field theory = Weinberg power counting



$$V_{1\pi}^{(0)}(\boldsymbol{p}',\boldsymbol{p}) = -\frac{g_A^2}{4f_\pi^2} \tau_1 \cdot \tau_2 \frac{\boldsymbol{\sigma}_1 \cdot \boldsymbol{q} \ \boldsymbol{\sigma}_2 \cdot \boldsymbol{q}}{q^2 + m_\pi^2}$$

$$V_{ct}^{(0)}(\boldsymbol{p}',\boldsymbol{p}) = C_S + C_T \ \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2$$

$$Central operator (no q dependence)$$

$$1\pi \text{ exchange (1PE)}$$
Pure contact term (CT)

4) Consistent construction of other operators (e.g. coupling to electroweak or WIMP probes)

Chiral effective field theory = interactions expansion



Ab initio (i.e. In medias res) quantum many-body problem



More effective approaches needed?

Consistent construction of other operators

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Specificities of atomic nuclei: mean field





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Bold = symmetry breaking (&restoration) single-reference methods



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Oxygen binding energies

• Oxygen chain: importance of **three-body forces** and **benchmark** case for ab initio calculations



✓ All methods yield consistent results within 2-3%
 ✓ 3N interaction mandatory

✓ Correct trend and drip-line location at N=16



- Neighbouring F & N chains
 Results are nicely consistent
- Interactions seem to work very satisfactorily

Extension to radii and mid-mas_s doubly closed-shell nuclei



[Entem, Machleidt 2003, Navrátil 2007, Roth et al. 2012]

Inconsistent Chiral N³LO 2N + N²LO 3N interactions

• Overbinding beyond ^AO that increases with mass

• Charge radii are consistently too small from ¹⁶O till ⁷⁸Ni

Successful	-ab initio description
	-of doubly closed-shell
	-ground states
	-up to A=78

Up to 20% error with data

Interaction uncertainties in doubly closed-shell nuclei



Many-body uncertainties in doubly closed-shell nuclei



• Fully converged calculations with respect basis size in mid mass nuclei

- Challenge however to go to heavier systems and refined many-body truncations
- Total many-body uncertainty in mid-mass nuclei ~2-3% dominated by NO2B approximation
- Uncertainty in mid-mass systems dominated by truncation in chiral EFT expansion

Binding energy in singly open-shell nuclei



Nuclear structure features addressed ab initio



Some challenges for ab initio theory

• More accurate descriptions

Next order in expansion, e.g. full T3, pert. T4
Next order in H, e.g. full 3NF and approx 4NF

• Enlarged portefolio of observables

- Low-lying E* in open-shell beyond sd
- Moments in open-shell beyond sd
- Giant resonances



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General conclusions

• Enormous progress of nuclear ab initio calculations in the last 10 years

- Much larger phenomenology can be put in connection with elementary nuclear forces
- \circ Nuclear forces themselves are explicitly rooted in QCD

 \circ Comparison to basic experimental observables can be made to day up to A ≈ 80

• Much further progress to be made

- Observables: electromagnatic moments and transitions, electroweak operators
- Nuclear interactions put to the test in mid-mass nuclei = current main bottleneck for progress
- Formal & numerical challenges to go to heavier nuclei/better accuracy/doubly open-shell nuclei
- Compute features of reactions (already some) and develop ab initio dynamics
- Evaluation and propagation of systematic errors of H

Recent/on-going cross-fertilization between nuclear physics and quantum chemistry

1) Symmetry breaking and restored methods: NP \rightarrow QC

BMBPT/BCC [A. Signoracci, T. Duguet, G. Hagen, G. R. Jansen, Phys. Rev. C91 (2015) 064320] [T. M. Henderson, G. E. Scuseria, J. Dukelsky, A. Signoracci, T. Duguet, Phys. Rev. C89 (2014) 054305]

 PBMBPT/PBCC
 [T. Duguet, J. Phys. G: Nucl. Part. Phys. 42 (2015) 025107]

 [Y. Qiu, T. M. Henderson, T. Duguet, G. E. Scuseria, Phys. Rev. C99 (2019) 044301]

2) Multi-reference methods: $QC \rightarrow NP$

 MRMBPT
 [Z. Rolik, A. Szabados, P. R. Surján, J. Chem. Phys. 119 (2003) 1922]

 [A. Tichai, E. Gebrerufael, K. Vobig, R. Roth, Phys. Lett. B786 (2018) 448]

3) In-medium similarity renormalization group method: NP \rightarrow QC

[K. Tsukiyama, S. K. Bogner, A. Schwenk, Phys. Rev. Lett. 106 (2011) 222502]

[A. Tichai, J. Toulouse, E. Giner , T. Duguet, work in progress]

4) Tensor-factorization techniques: $QC \rightarrow NP$

[R. Schutski, J. Zhao, T. M. Henderson, G. E. Scuseria, J. Chem. Phys. 147 (2017) 184113]

[A. Tichai, R. Schutski, G. E. Scuseria, T. Duguet, Phys. Rev. C99 (2019) 034320]

5) Importance truncation techniques for, e.g., (P)(B)CC: $QC \rightarrow NP$

[J. E. Deustua, J. Shen, P. Piecuch, PRL 119 (2017) 223003]

[A. Tichai, J. Ripoche, T. Duguet, Eur. Phys. J. A55 (2019) 90]
Collaborators on ab initio many-body calculations



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Elementary facts and « big » questions about nuclei

• 252 **stable** isotopes, ~3100 synthesized in the lab • **How many** bound (w.r.t strong force) nuclei exist; 9000? 120 Stable nuclei Unstable nuclei N = 126Predicted bound nuclei 100 Z = 8280 Proton number (Z) N = Z symmetry 60 Z = 5040 Z = 2820 = 20 [figure from Bazin 2012] 180 60 100 120 140 160 20 40 80 Neutron number (N)

Updated in 2019 to Z=9 (22 neutrons) and Z=10 (24 neutrons)

- Neutron **drip-line** known up to Z-8 (16 neutrons)
- \circ Where is the neutron drip-line beyond Z=10?

Oganesson (118Og) added to Mendeleïev table in 2016

- **Heaviest** synthesized element Z=118
- Heaviest possible element?

• Enhanced stability near Z=120?126?

Modes of instability (α, β, γ decays, fission)
Are there more exotic decay modes?
Ex: ν-less 2β decay = test of standard model
Ex: 2p decay beyond the proton drip line

Elements up to Fe produced in stellar fusion
How have heavier elements been produced?
r-process nucleosynthesis in neutron star mergers

Over-stable "magic" nuclei (2, 8, 20, 28, 50, 82, ...)
Are magic numbers the same for unstable nuclei?

Single-reference expansion many-body methods

Nuclear Hamiltonian

 $H = T + V^{2N} + W^{3N}$

Symmetry group U(1) dealt with today [H,S] = 0 where $S \equiv A, J^2, J_z...$



Non-degenerate Good starting point No**Dedegenara**te I**Rpopprestaratitig g pioii**nt A-body eigenvalue problem

 $H|\Psi_0^{\rm S}\rangle = E_0^{\rm S}|\Psi_0^{\rm S}\rangle$ N^A cost where N = dim H₁

Many-body expansion



Wave operator Reference state

Accounts for « weak/dynamical » correlations
 Expand as a series (MBPT, CC...) + truncate = N^p cost

 $[H'_0, S] \neq 0$ $[H'_1, S] \neq 0$

 $H = H'_{0} + H'_{1}$ $|\Psi_{0}^{S}\rangle = U(\infty)|\Phi_{0}\rangle$ More general reference state

Symmetry breaking

Accounts for "strong/non-dynamical" correlations
 Expand (BMBPT, BCC...) + truncate = N^p cost

- 1) Truncated series breaks symmetry
- 2) Exact symmetry must eventually be restored

First ab initio calculations

1990's: Green function Monte Carlo approach [Carlson, Pieper, Wiringa, Schiavilla,...]
 MC techniques to sample many-body wave function in coordinate, isospin and spin space
 2000's: No-core shell model approach (i.e. full CI) [Vary, Barrett, Navratil, Ormand...]
 Diagonalisation of the Hamiltonian in a finite-dimensional space



Nuclei simulated from "scratch"! Closed the gap between elementary inter-nucleon interactions and properties of nuclei

[Pieper & Wiringa 2001]

X Computational effort increases exponentially/factorially with nucleon numberX Necessity of treating three-nucleon forces makes it more severe

→ Approach limited to light nuclei (~A≤12)

Chiral EFT hamiltonians

N3LO (~2010)

[Entem & Machleidt 2003, Navrátil 2007, Roth et al. 2012]

- First generation of ChEFT interactions (N³LO 2° , LO 3N)
- Follows traditional ab initio strategy (fit G y sector on X-body data)
- Successful in light nuclei, but strong Jverbinding and too small radii for heavier systems

• NNLOsat (2015)

[Ekström et al. 2015]

- Development prompted by inability to reproduce radii beyond light nuclei
- Data from not-so-light nuclei (*A*=14-25) included in fit + Non-local 3NF regulator
- Good BE and radii in mid-mass but two- and few-few-body systems slightly deteriorated

• N3LOInl (2018)

- Back to standard ab initio strategy but with important intation of non-local regulators
- Correct description of two- and few-being Lems
- BE and radii of mid-mass systems much improved compared to N3LO

[Entem & Machleidt 2003, Navrátil 2018]

Sources of uncertainty



• Many-body truncation typically 2-3%



• Difference with data up to 10-15% in Ca-Ni region with N3LO



Largest uncertainty from input Hamiltonian

→ Improved Hamiltonians needed

Charge radii in medium-mass nuclei



• Newly developed Hamiltonians improves the situation



Charge radii in medium-mass nuclei



Charge radii in medium-mass nuclei



Output: Content of the stringent tests of nuclear interactions via ab initio calculations of mid-mass chains

Ab initio emergence of N=20 and N=28 magic numbers



Spectra of Fluorine isotopes

• Excitation spectra of (neutron-rich) ^{19,23,25,26}F from *ab initio* sd shell model

N3LO (~2010)

 \circ Hybrid method = ab initio shell model (core ¹⁶O and valence space H from IMSRG)



✓ Very satisfactory account of experimental data

✓ 3N interaction mandatory for correct density of states and ordering

✓ As good as best sd shell empirical USDB interaction (i.e. traditional shell model)

Confrontation with spectroscopic data in sd nuclei can now be based on ab initio scheme!

[Stroberg et al. 2016]

Spectra of K isotopes



Potential bubble nucleus ³⁴Si



• Excellent agreement with experimental charge distribution of ³⁶S [Rychel et al. 1983]

• Charge density of ³⁴Si is predicted to display a marked depletion in the center

Charge form factor

• Charge form factor measured in (e,e) experiments sensitive to bubble structure? **PWBA** $F(q) = \int d\vec{r} \rho_{\rm ch}(r) e^{-i\vec{q}\cdot\vec{r}}$ with momentum transfer $q = 2p\sin\theta/2$ --- ³⁴Si [ADC(2)] 100 LOI accepted SCGF ³⁴Si [ADC(3)] 10 $\langle r^2 \rangle^{1/2}$ [fm] Nuclei I^{π} μ [nm] Q[b][Duguet *et al.* 2017] $T_{1/2}$ --- ³⁶S [ADC(2)] 0^{+} ^{24}Si 140 ms25Si $5/2^+$ 220 ms ³⁶S [ADC(3)] 0.1 ²⁶Si 2.2 s 0^{+} **NNLO**_{sat} $|F(\theta)|^2$ 27Si $4.1 \mathrm{s}$ $5/2^+$ (-)0.8554(4) (+)0.060(13)10-2 ²⁸Si 0^{+} 3.106(30)stable ²⁹Si $1/2^{+}$ -0.55529(3)3.079(21)stable 10-3 ³⁰Si 0^{+} stable 3.193(13)³¹Si $3/2^{+}$ 10-4 157.3 m ^{32}Si 153 v 0^{+} 10-5 ³³Si $(3/2)^+$ (+)1.21(3)61 s E = 300 MeV³⁴Si 2.8 s 0^{+} 10-6 ³⁵Si 0.8 s $(7/2)^{-}$ (-)1.638(4)10-7 60 80 100 20 40 120 140 160 0 θ [deg]

• Central depletion reflects in larger $|F(\theta)|^2$ for angles 60°< θ <90° and shifted 2nd minimum by 20°

- Visible in future electron scattering experiments if enough luminosity (10²⁹ cm⁻²s⁻¹ for 2nd minimun)
- Correlation between F_{ch} and $\langle r^2 \rangle_{ch}^{1/2} ({}^{36}S) \langle r^2 \rangle_{ch}^{1/2} ({}^{34}Si)$ identified

■Measurement of $\delta < r^2 >_{ch}^{1/2}$ (^ASi) from high-resolution laser spectroscopy@NSCL (R. Garcia-Ruiz)

Addition and removal nucleon spectra



• **Good agreement** for one-neutron addition to ³⁵Si and ³⁷Si $(1/2^{-} \text{ state in } {}^{35}\text{Si} \text{ needs continuum})$ • Much less good for one-proton removal; ³³Al on the edge of island of inversion: challenging!

• Correct reduction of splitting E_{1/2}⁻ - E_{3/2}⁻ from ³⁷S to ³⁵Si

Such a sudden reduction of 50% is unique Any correlation with the bubble? Yes!

$E_{1/2^{-}} - E_{3/2^{-}}$	^{37}S	³⁵ Si	$^{37}S \rightarrow ^{35}Si$
SCGF	2.18	1.16	-1.02 (-47%)
(d,p)	1.99	0.91	-1.08 (-54%)

Electromagnetic response

• Photodisintegration cross section of ⁴⁰Ca

DysADC3 RPA

 40 Ca

NNLO_{sat}

••• Ahrens (1975)

 N_{max} =13, $\hbar\omega$ =20 MeV

[Raimoni and Barbieri 2019]

 $\sigma(\mathbf{E}_X)$ [mb]

SCGF

 $\tilde{\mathbf{E}}_X$ [MeV]



Dipole response function

$$R(E) \equiv \sum_k |\langle \Psi_k | Q_{1m}^{T=1} | \Psi_0 \rangle | \delta(E_k - E_0 - E)$$

Electric dipole operator

$$Q_{1m}^{T=1} \equiv \frac{N}{N+Z} \sum_{p=1}^{Z} r_p Y_{1m}(\theta_p, \phi_p) - \frac{Z}{N+Z} \sum_{n=1}^{N} r_n Y_{1m}(\theta_n, \phi_n)$$

Giant and pygmy resonances accessible up to ^ANi Many-body correlations crucial for quantitative description



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So what about observables from laser spectrocopy?

• Charge radii via isotopic shifts

- \circ Tremendously useful to tune bulk properties of nuclear interactions
- Now systematically computed for even-even closed and (singly) open-shell nuclei
- \circ Entertain interesting correlations with other observables, e.g. α_D , F_{ch} ...

• Nuclear spins via atomic hyperfine structure

- Basic check of nuclear structure evolution
- Require the computation of odd-even or odd-odd ground-states/isomeric states
- Systematic comparison with available data could be useful

• Ground-state electromagnetic moments via atomic hyperfine structure

- Detailed probe of nuclear structure evolution (« shell structure » and « shell occupancies »)
- \circ Require the computation of odd-even or odd-odd ground-states
- **•** Require the computation of non-trivial operators

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Consistent operators in chiral effective field theory

✓ Nuclear electromagnetic charge/current operators (= time/vector part of four-vector current j^µ)

$$\rho(\vec{q}) = \sum_{i} \rho_{i}(\vec{q}) + \sum_{i < j} \rho_{ij}(\vec{q}) + \sum_{i < j < k} \rho_{ijk}(\vec{q}) + \dots$$

$$(a) One-body (i.e. standard) operator$$

$$(b) Two-body meson-exchange currents (MECs)$$

$$(c) Three-body meson-exchange currents$$

• Operators are built from EFT expansion by coupling nuclear current to external e.m. fields

• Consistent nuclear e.m. operators and nuclear forces

• Satisfy the continuity equation $\vec{q} \cdot \vec{j}(\vec{q}) = [H, \rho(\vec{q})]$ ollowing from gauge invariance

• Derived via two different version of time-ordered perturbation theory

- Standard time-ordered perturbation theory / Jlab-Pisa group [Pastore et al. 2008, 2009, 2011, 2013]
- Method of unitary transformation / Bochum-Bonn group [Kolling et al. 2009, 2011]

Proper renormalization achieved in this case

Electromagnetic current operator



Electromagnetic charge operator



Relation to observables from laser spectroscopy

• Longitudinal and transverse form factors for elastic and inelastic scattering

$$\begin{split} F_{L}^{2}(q) &= \frac{1}{2J_{i}+1} \sum_{J=0}^{\infty} |\langle \Psi_{f}^{J_{f}} \overline{[T_{J}^{C}(q)]} \Psi_{i}^{J_{i}} \rangle|^{2} \\ F_{T}^{2}(q) &= \frac{1}{2J_{i}+1} \sum_{J=0}^{\infty} |\langle \Psi_{f}^{J_{f}} \overline{[T_{J}^{M}(q)]} |\Psi_{i}^{J_{i}} \rangle|^{2} + |\langle \Psi_{f}^{J_{f}} \overline{[T_{J}^{E}(q)]} \Psi_{i}^{J_{i}} \rangle|^{2} \\ \end{split}$$

$$\begin{aligned} \mathbf{T}^{\mathsf{C}_{\mathsf{J}}} \leftarrow \text{multipole expansion of } \mathsf{p} \end{aligned}$$

- Connection to static moments
 - → Elastic scattering on ground-state: $J_i = J_f = J_0$ → Static limit: q = 0 $T_1^C(0) \propto Q$ $T_1^M(0) \propto \mu$
- Form of standard one-body, i.e. LO(IA), operators

$$Q^{\mathrm{IA}} = e \sum_i e_i(0) \, r_i^2 \, Y_{20}(\theta_i,\phi_i)$$

• Static magnetic dipole operator

$$\mu^{\text{IA}} = \sum_{i} e_i(0) \vec{L}_i + \mu_i(0) \vec{\sigma}_i$$

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Magnetic dipole moment in s and p shell nuclei





- Decomposition of one-body IA
 - Proton's convection small vs spin magnetization
 - Driven by valence nucleon in odd-even
 - Driven by n-p or 3He-p cluster in odd-odd

Elastic form factors in s and p shell nuclei

• Elastic charge (longitudinal) and magnetic (transverse) form factors from ²H to ¹²C



Ex: Quadrupole electric form factor in ²H
 Hybrid and (semi-consistent) χ-EFT calculations
 Charge operator at LO (IA) and N³LO
 Band from 500 MeV < Λ < 600 MeV

Results

- \circ G_Q(0) = M²_d Q_d (here in fit of NN)
- \circ LO(IA) sufficient up to q~3 fm-1
- \circ Nucleonic form factors mandatory beyond 1.5 $\rm fm^{-1}$
- \circ Excellent result up to q ~ 4 fm⁻¹ in all cases
- \circ $\chi\text{-EFT}$ with N^3LO MEC excellent up to q ~ 8 fm^{-1}

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Moments in Ca isotopes

• Empirical/ab initio (IMSRG) shell-model calculations of magnetic dipole/electric quadrupole momer

- ^{47,49,51}Ca via high-resolution collinear laser spectroscopy COLLAPS @ ISOLDE [Garcia Ruiz et al. 2015]
- ³⁷Ca via collinear laser spectroscopy BECOLA @ NSCL [Klose et al. 2019]



Operators

- \circ Pure one-body \leftrightarrow No explicit MEC
- Bare spin and orbital g factors for magnetic moment
- \circ Effective charges: $e_n = 0.5e$ and $e_p = 1.5e$

Magnetic moment

- \circ ⁴⁰Ca core broken in ^{41,43,45}Ca
- Good reproduction from ab initio in ^{47,49,51}Ca



Significant breaking of N=32 magic number

• Quadrupole moment

• Excellent agreement for ab initio in all isotopes

 \circ No apparent need of orbital-dependent e_n and / or e_p

Next: MEC and consistently-transformed operators to valence space

Huge diversity of nuclear phenomena



Ab initio (i.e. In medias res) quantum many-body problem



Ab initio vs effective approach



Matter of choice

- Which properties we aim at and which level of accuracy are we seeking?
- Applicability throughout the nuclear chart? \rightarrow Universal/global vs local description
- Wish to connect to underlying theory of strong force or wish to focus on describing data?
- Predictive power? → Estimate of theoretical error

Chiral effective field theory = basic considerations



Similarities and differences with quantum chemistry


Hamiltonian

When working in Fock space

Nuclear Hamiltonian

Grand potential

 $\Omega \equiv H - \lambda A \checkmark$

Particle number

$$H = \frac{1}{(1!)^2} \sum_{pq} t_{pq} c_p^{\dagger} c_q$$

+ $\frac{1}{(2!)^2} \sum_{pqrs} \overline{v}_{pqrs} c_p^{\dagger} c_q^{\dagger} c_s c_r$
+ $\frac{1}{(3!)^2} \sum_{pqrstu} \overline{w}_{pqrstu} c_p^{\dagger} c_q^{\dagger} c_r^{\dagger} c_u c_t c_s$

$$A \equiv \sum_{p} c_{p}^{\dagger} c_{p}$$

Genuine 3N interaction / six-legs vertex

k-body force ↓ Mode-2k tensor

Basis representation dim N

Storage cost N^{2k}



Problematic to handle 3N interactions in mid-mass nuclei

Controls the average particle number in the system

Chemical potential

Slater determinant reference state and normal ordering



$$+ \frac{1}{3!3!} \sum_{l_1 l_2 l_3 l_4 l_5 l_6} \Lambda^{33}_{l_1 l_2 l_3 l_4 l_5 l_6} : c_{l_1}^{\dagger} c_{l_2}^{\dagger} c_{l_3}^{\dagger} c_{l_6} c_{l_5} c_{l_4}$$

Six-index tensor Too expensive to handle

NO2B approximation
 1-3% error in closed shell
 [R. Roth et al., PRL 109 (2012) 052501]

Effective 2-body operators Captures essential of 3-body Many-body method with 2-body

Bogoliubov reference state and normal ordering

Bogoliubov reference state

Breaks U(1) symmetry

$$\beta_{k} = \sum_{p} U_{pk}^{*} c_{p} + V_{pk}^{*} c_{p}^{\dagger} \qquad |\Phi\rangle \equiv C \prod_{k} \beta_{k} |0\rangle \qquad A|\Phi\rangle \neq A|\Phi\rangle$$

$$\beta_{k}^{\dagger} = \sum_{p} U_{pk} c_{p}^{\dagger} + V_{pk} c_{p} \qquad \beta_{k} |\Phi\rangle = 0 \quad \forall k \qquad \text{Vacuum state} \\ \text{Reduces to SD in } \mathcal{H}_{A} \text{ for closed-shell}$$

Normal ordering via Wick's theorem in quasi-particle basis

$$H^{ij} \text{ matrix elements function of}$$

$$H \equiv \sum_{n=0}^{3} \sum_{i+j=2n} \frac{1}{i!j!} \sum_{l_1...l_{i+j}} H^{ij}_{l_1...l_{i+j}} \beta^{\dagger}_{k_1} \dots \beta^{\dagger}_{k_i} \beta_{k_{i+j}} \dots \beta_{k_{i+1}}$$

$$= H^{00} + [H^{20} + H^{11} + H^{02}] + [H^{40} + H^{31} + H^{22} + H^{13} + H^{04}] + \sum_{i+j=6} H^{ij}$$

$$= \sum_{n=0}^{2} H^{[2n]} + H^{[6]} \quad 6\text{-qp operators}$$
Six-index tensors
Too expensive to handle
NO2B approximation
1-3% error in closed shell
[Roth et al., PRL 109 (2012) 052501]
$$H^{ij} \text{ matrix elements function of}$$

$$PNO2B approximation$$

$$PNO2B approximation$$

$$Particle-number conserving[Ripoche, Tichai, Duguet, arXiv:1908.00765]$$

Electron scattering off nuclei

- Electrons constitute an optimal probe to study atomic nuclei
 Point-like → excellent spatial resolution
 EM weak and theoretically well constrained
- Accélérateur Linéaire @ Saclay (ALS)
 - Electron accelerator (1969-1990)
 - Refined data on tens of stable nuclei









[Frois et al. 1977]

➡ Electron scattering off unstable nuclei?

- \circ Challenge for the future
- \circ First physics experiments in 2017 with SCRIT @ RIKEN

Guidance for improved nuclear many-body Hamiltonians

Nuclear lattice calculations of 86 even-even nuclei up to A=48 and pure neutron matter

[Lu et al. 2018]

¤ Leading-order pion-less EFT SU(4)-invariant with 2N and 3N interactions



Error < 4.5% on BE in ¹⁶O and < 8.0% on R_c in ³H

Effective range r₀ averaged over ¹S₀ and ³S₁ S-wave scattering length a₀ averaged over ¹S₀ and ³S₁ B(³H) + set of mid-mass nuclei



SU(4)-invariant LO very satisfactory for large A
 Satisfatory pure neutron matter + volume/surface energy coefficients
 Corrections from spin&isospin dependent terms

Coulomb effect beneficial

Novel many-body formalisms

No real free lunch but still look for best compromise

 Versatility (nuclei and / or states / observables
 Accuracy

✓ CPU cost

[Duguet, Signoracci 2016]

• Optimal many-body method for open-shell nuclei: **Bogoliubov many-body perturbation theory**

→ Code for automated generation of many-body diagrams [*Arthuis et al.* 2018]



 \rightarrow 2-3% agreement of all methods with exact results (IT-NCSM)

→ Different truncation schemes yield consistent description of open-shell nuclei

BMBPT optimal to systematically test next generation of Chiral EFT nuclear Hamiltonians